



Cornell University
Laboratory for Elementary-Particle Physics



ECLLOUD Calculations of Coherent Tune Shifts for the 6.4 km ILC Damping Ring

- Including an estimate of the rediffused SEY component contribution -

*With reference to Theo Demma's talks of 15 Dec 2009 and 10 Jan 2010
and Mauro Pivi's talk of 3 Feb 2010*

Jim Crittenden

Cornell Laboratory for Accelerator-Based Sciences and Education

ILC Damping Ring Electron Cloud Working Group Meeting

23 February 2010





Theo's ECLLOUD input parameters for the 6.4 km ILC damping ring

Bunch population	N_b	2.1×10^{10}
Number of bunches	N_b	45 x 4 trains
Bunch gap	N_{gap}	15
Bunch spacing	$L_{sep}[m]$	1.8
Bunch length	$\sigma_z[mm]$	6
Bunch horizontal size	$\sigma_x[mm]$	0.26
Bunch vertical size	$\sigma_y[mm]$	0.006
Photoelectron Yield	Y	0.1
Photon rate ($e^-/e^+/m$)	dn_γ/ds	0.204
Antechamber protection	η	90%; 97%
Photon Reflectivity	R	20%; 50%
Max. Secondary Emission Yield	δ_{max}	1.2
Energy at Max. SEY	$E_m[eV]$	300
SEY model	Cimino-Collins ($\delta(0)=0.5$)	

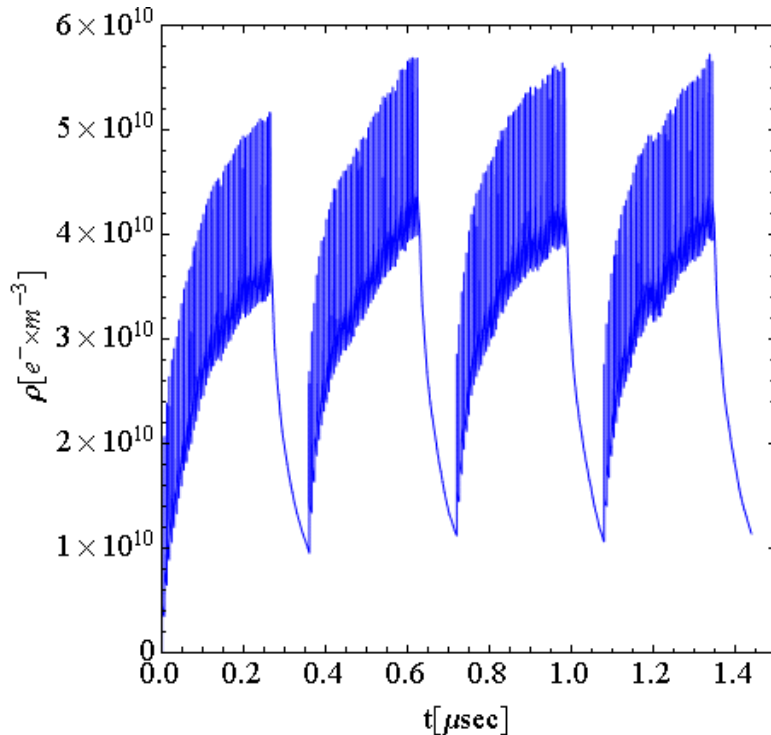
*CesrTA ECLLOUD development version used for the case $\eta=90\%$, $R=20\%$.
Primary p.e. model for this version differs in the generated 2D angular distribution.
Also, the SEY model for true secondaries has a slightly different energy dependence (à la POSINST).*



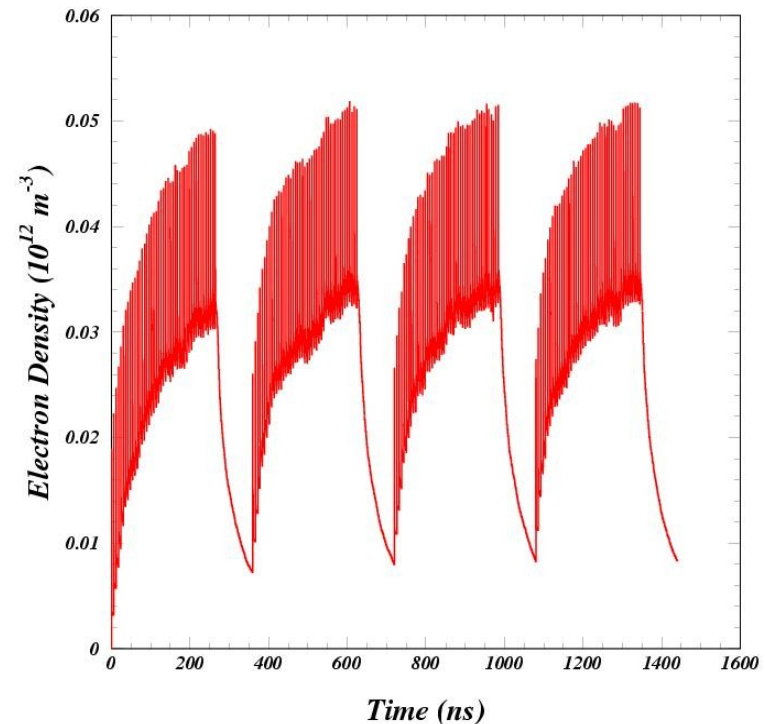
Theo's presentation of 15 Dec 2009

By=0.27 T; R=20%

$\eta=90\%$ $\eta=97\%$



*Calculation repeated with CestrTA
development ECLLOUD version for
By = 0.27 T, $\eta = 90\%$, R = 20%*



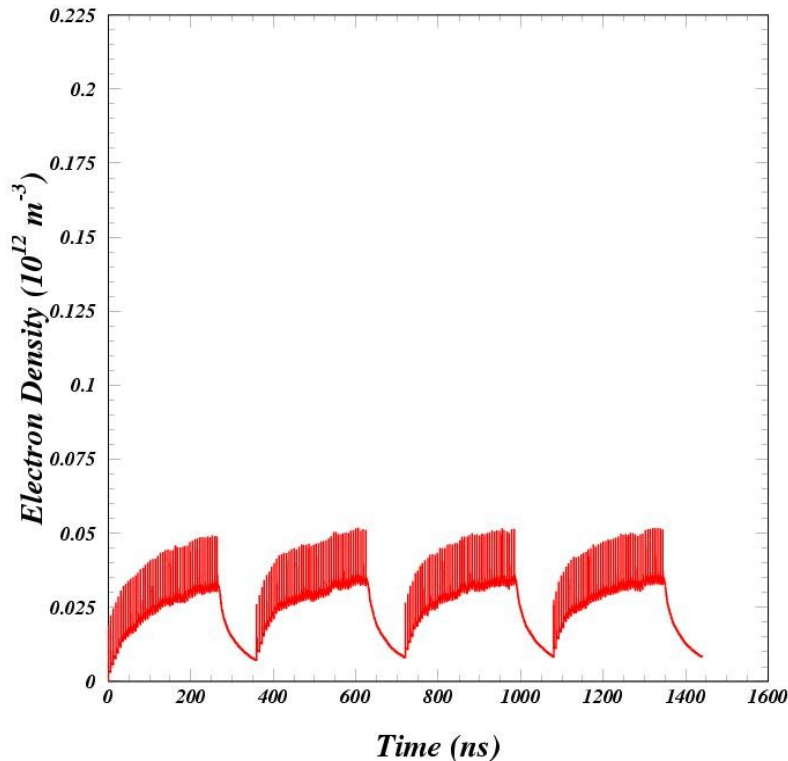
The two calculations agree reasonably well.

ECLLOUD default primary angular distribution produces more grazing incidence, thus higher secondary yield.

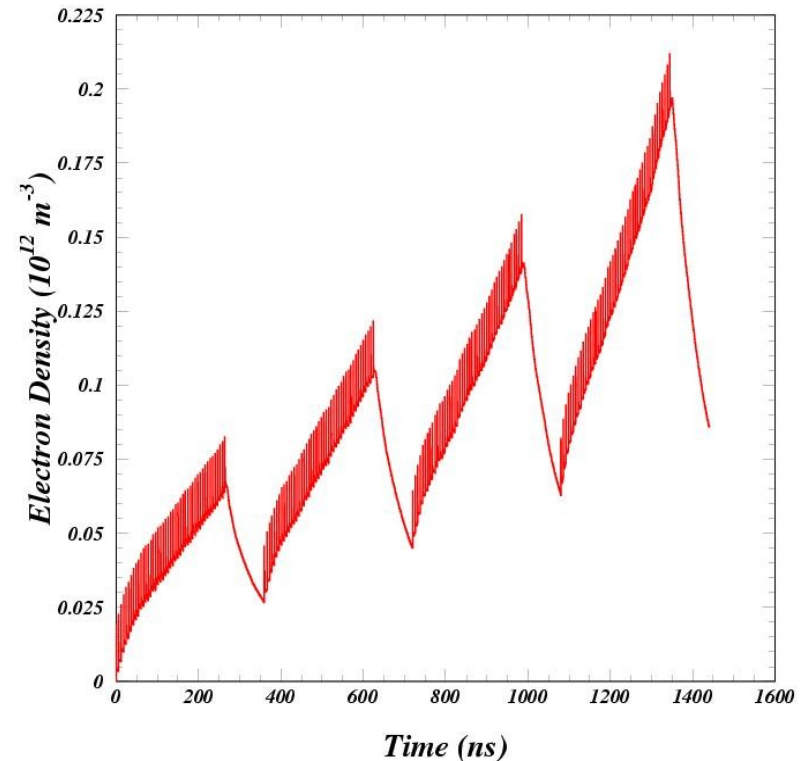


Effect of the rediffused SEY component on the beampipe-averaged cloud density buildup

Original ECLLOUD SEY model excludes the rediffused SEY component



Cloud buildup including the rediffused SEY component used in the CsrTA tune shift model ($\delta = 0.2$)

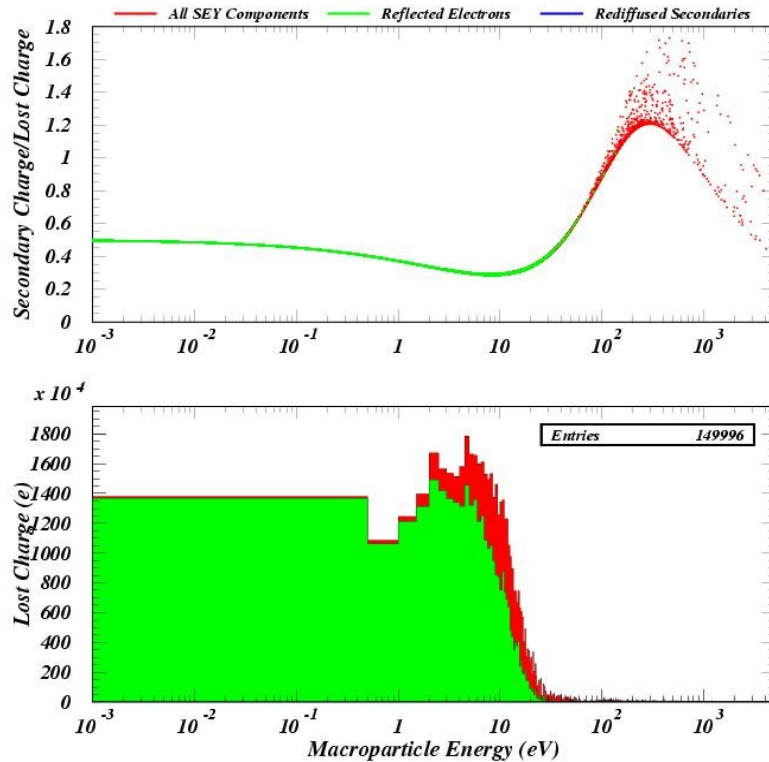


The rediffused component has a particularly strong effect in dipole magnets, where the cloud particles are trapped on vertical field lines. The kinematics of the rediffused secondaries raise the average energy of the cloud particles hitting the wall and therefore the secondary yield (in some circumstances).

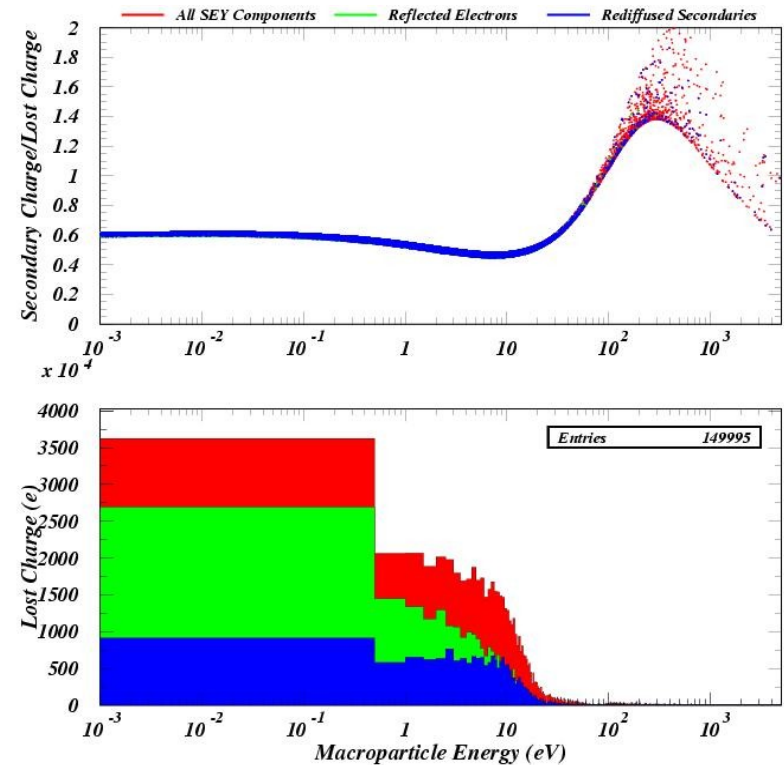


Cloud buildup, snapshot, profile and SEY plots available at <http://www.lepp.cornell.edu/~critten/ilcdr/23feb10>. Also for an electron beam.

Rediffused SEY component excluded



Rediffused SEY component included

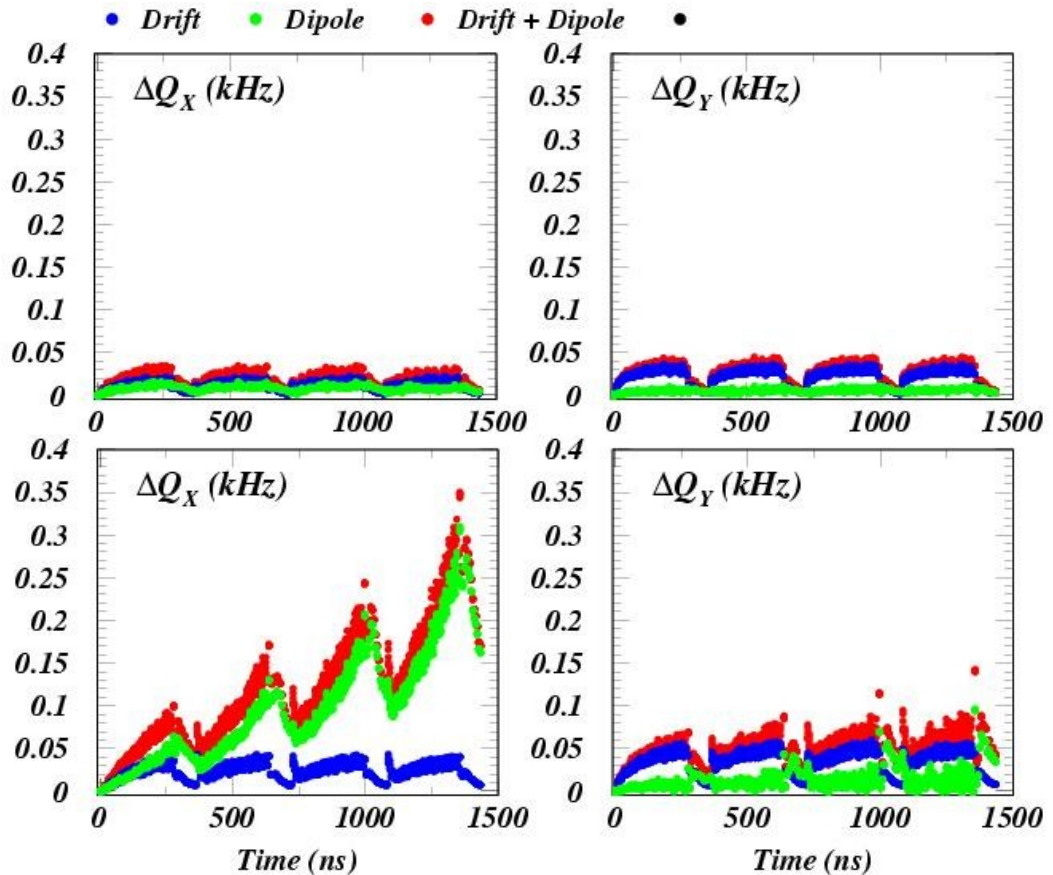


NB: The true secondary yield was not scaled down by the rediffused yield, as POSINST does.



No rediffused SEY component

*Rediffused SEY component
included*



*Assumed 80% dipole and 20% drift. Assumed average beta functions of 10 m.
Better values can be obtained from Kiran's synchrotron radiation analysis of the lattice file,
which will also provide detailed element- and beta-averaged photon flux values.*

Also need to decide on more realistic values for the rediffused SEY yield.