

### e-Cloud Theoretical Studies

Levi Schächter

# Outline

- Photo-Electrons
- Dynamics of Electrons in e-Cloud
- e-Cloud Spatial Distribution
- Wake in the e-Cloud
- Experimental Implications

	$E_{\rm th}[eV]$	$E_k = 2[GeV]$	$E_k = 5[GeV]$
N <sub>ph</sub> / N <sub>e</sub>	0	259.4	648.5
	4	177 (68%)	565 (87%)
	300	13.8(5%)	314 (48%)
	0	177[kV]	0.690[MV]
$V_{ m ph}$	4	176[kV]	0.690[MV]
	300	6.6[kV] (4%)	0.667[MV] (97%)
Peak of SE (Al)		Work function (Al)	3

# Photo-Electrons Yield

Photo-electrons spectrum is a "convolution" with the photon's spectrum

$$\tilde{N}_{\rm pe}(E) \simeq \int_{E_{\rm w}}^{\infty} dE' \delta_{\rm pe}(E|E') \left[ N_{\rm ph} f_{\rm ph}(E') \right]$$

Similar to the photons we are interested on :

$$\frac{\overline{N}_{pe}(E > E_{th})}{N_{e}} = 49.6 \times E_{k}[GeV] \int_{E_{th}}^{\infty} dE \int_{E_{w}}^{\infty} dE' \frac{1}{E_{cr}} \delta_{pe}(E|E') \int_{2E'/E_{cr}}^{\infty} dx K_{5/3}(x)$$
$$V_{pe}(E > E_{th}) = 2.193 \times 10^{5} \frac{E_{k}^{4}(GeV)}{\rho[m]} \int_{E_{th}}^{\infty} dE \frac{E}{E_{cr}} \int_{E_{w}}^{\infty} dE' \delta_{pe}(E|E') \int_{2E'/E_{cr}}^{\infty} dx K_{5/3}(x)$$

#### Photo-Electrons – Model for the Yield

**1.** According to Grobner et.al [J. Vac. Sci. Technolo. A 7 (2) 1989 p, 223-9 "p.224 (Figure 2)] the yield of extracting a photo-electron from an aluminium surface

$$\delta_{\rm pe}(E|E') \propto Y(E') = \left(\frac{E_0}{E'}\right)^{\nu} \qquad \text{Typically:} \begin{cases} E_0 = 2.316[eV] \\ \nu = 0.825 \end{cases}$$

2. The energy of the photo-electron should not exceed the energy of the incident photon and we further assume that the electrons' energy is uniformly distributed between zero and the energy of incoming photon

$$\delta_{\mathrm{pe}}(E|E') \propto \frac{1}{E'} [h(E) - h(E - E')]$$



### Photo-Electrons – Model for the Yield

3. No contribution from photons of energy lower than the work function

 $\delta_{\rm pe}(E|E') \propto h(E'-E_{\rm w})$ 

4. Not all the photons are absorbed **Reflection coefficient**  $\rho(E')$  $\delta_{pe}(E|E') \propto \left[1 - \left|\rho_{eff}(E')\right|^{2}\right]$ 

"Handbook of X-rays", edited by E.F. Kaelble, McGraw-Hill, New York, 1967, p. 48-2  $\left|\rho_{\rm eff}\left(E\right)\right|^{2} \approx \frac{1.881}{E} \exp\left(-1.2486 \times 10^{-4} E^{2}\right)$ 

Virtually all photons with energy higher than 100eV are absorbed at the first impact!!

$$\delta_{\rm pe}\left(E|E'\right) \simeq Y(E')\frac{1}{E'}\left[h(E) - h(E-E')\right]h\left(E' - E_{\rm w}\right)\left[1 - \left|\rho_{\rm eff}\left(E'\right)\right|^2\right]$$

# 

### Photo-Electrons –Preliminary Conclusions

Comparable

	$E_{\rm th}[eV]$	$E_k = 2[GeV]$		$E_k = 5[GeV]$	
		$\rho_{\rm eff}\left(E\right) = 0$	$\rho_{\rm eff}\left(E\right) \neq 0$	$\rho_{\rm eff}\left(E\right) = 0$	$\rho_{\rm eff}\left(E\right) \neq 0$
$N_{\rm pe}$ / $N_e$	0	25.8	20.9	32.9	27.6
	300 <	0.047(0.18%)	0.047(0.22%)	1.1(3.3%)	1.1(4%)
$V_{ m pe}$	0	354.4[V]	332.9[V]	1.64[kV]	1.61[kV]
	300	19.8[V](5.6%)	19.8[V](6%)	0.97[kV](59%)	0.97[kV](60%)
$\overline{\delta}_{ m pe}$	0	0.1(0)	0.8(-1)	0.5(-1)	0.4(-1)
	300	1.8(-4)	1.8(-4)	1.7(-4)	1.7(-4)

One order of magnitude difference

Reflections (Al) may reduce the number of photo-electrons by about 20%.

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 $\frac{\partial \tilde{N}_{\rm ec}\left(t,E\right)}{\partial t} = \frac{1}{\tau\left(E\right)} \left[ -\tilde{N}_{\rm ec}\left(t,E\right) + \tilde{N}_{\rm se}\left(t,E\right) \right] + \tilde{N}_{\rm pe}\left(E\right) \sum_{n=1}^{N_b} \delta\left(t-T_n\right)$ Spectrum of SE Life-time Spectrum of PE **Bunch** separation is much longer than .. or Ionization electrons the duration of single bunch 9

# Dynamics of Electrons in e-Cloud $\tau(E) = ??$



$$\frac{d^2 y}{dt^2} = -\frac{e}{m} E_y = \frac{e^2 n_0}{m\varepsilon_0} y = \omega_p^2 y$$
$$Y\left(\tau \equiv \omega_p t, \xi \equiv \frac{V_0}{\omega_p D}\right) \equiv \frac{y(t)}{D} = \cosh(\tau) - \xi \sinh(\tau)$$



$$\tau(E) \simeq \frac{\pi}{\omega_p} \begin{cases} \frac{\sqrt{E_0 / E}}{2 - (E_0 / E)^2} & E > E_0 \\ \sqrt{E / E_0} & E < E_0 \end{cases}$$

$$E_0 = \frac{1}{2}m\omega_p^2 a_y^2$$



$$E = E_0 = 100[eV], \ D_y = 2.5[cm]$$

$$n_{\max}[m^{-3}] = 1.1 \times 10^{12} \frac{E[eV]}{a_y^2[cm]} = 1.76 \times 10^{13}$$

$$\overline{\tau}^{(\max)}[n \sec] = 52.93 \frac{a_y[cm]}{\sqrt{E[eV]}} = 13.2$$

Note that the life-time of the electron in the cloud may be comparable to the bunch separation !!!

# Dynamics of the e-Cloud

Let us consider the dynamics w/o SE

$$\frac{\partial \tilde{N}_{ec}(t,E)}{\partial t} = \frac{1}{\tau(E)} \left[ -\tilde{N}_{ec}(t,E) + \tilde{N}_{se}(t,E) \right] + \tilde{N}_{pe}(E) \sum_{n=1}^{N_b} \delta(t-T_n)$$

$$\tilde{N}_{ec}(t,E) = \tilde{N}_{pe}(E) \sum_{n=1}^{N_b} h(t-T_n) \exp\left[ -\frac{t-T_n}{\tau(E,\bar{n}_{ec})} \right]$$
Both build-up  
and steady-state  
depend on the  
ratio  $T/\tau$ 

$$y_{n+1} = (y_n + \Delta) \exp(-T/\tau)$$
 $t/T$ 
 $t/T$ 

In practice, we need to integrate over the entire spectrum

$$\overline{n}_{\rm ec}(t) \equiv \frac{1}{\left(D_x D_y\right) \left(2\pi\rho\right)} \int_0^\infty dE \widetilde{N}_{\rm ec}(t, E)$$
$$= \frac{\overline{N}_{\rm pe}}{\left(D_x D_y\right) \left(2\pi\rho\right)} \sum_{n=1}^{N_b} h\left(t - T_n\right) \int_0^\infty dE f_{\rm pe}(E) \exp\left[-\frac{t - T_n}{\tau(E, \overline{n}_{\rm ec})}\right]$$

For the parameters involved, we found that defining the average life-time  $\overline{\tau}(\overline{n}_{ec}) \equiv \int dE f_{pe}(E) \tau(E, \overline{n}_{ec}) \qquad ^{2}$ 

is convenient since

$$\overline{n}_{\rm ec}(t) = \frac{\overline{N}_{\rm pe}}{\left(D_x D_y\right) \left(2\pi\rho\right)} \sum_{n=1}^{N_b} h\left(t - T_n\right) \exp\left[-\frac{t - T_n}{\overline{\tau}\left(\overline{n}_{\rm ec}\right)}\right]$$

No significant difference between the two



exact

And now with secondary emission included:  $\tilde{N}_{se}(t,E) = \int_{0}^{\infty} dE' \delta_{se}(E|E') \tilde{N}_{ec}(t,E')$ 

**Assumption:** Because of the vertical magnetic field (bend or wiggler) the electrons gyrate along the magnetic field line consequently, both incoming and outgoing electrons are assumed to be moving in the vertical direction only.

Without significant loss of generality, based on the experimental data, it is reasonable to further assume that the energy of the secondary electrons is uniformly distributed between zero and that of the incident electron

$$\delta_{\rm se}(E|E') \simeq [h(E) - h(E - E')] \frac{\tilde{\delta}_{\rm se}(E')}{E'}$$

Consequently,

$$\frac{\partial \tilde{N}_{\rm ec}(t,E)}{\partial t} = \frac{-1}{\tau(E)} \left[ \tilde{N}_{\rm ec}(t,E) - \int_{E}^{\infty} dE' \frac{\tilde{\delta}_{\rm se}(E')}{E'} \tilde{N}_{\rm ec}(t,E') \right] + \tilde{N}_{\rm pe}(E) \sum_{n=1}^{N_b} \delta(t-T_n)$$
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*A realistic model for secondary emission [*J.R.M Vaughan, "A new Formula for Secondary Emission Yield", IEEE Transactions on Electron Devices Vol 36 (9) p.1963 (1989). Actually a variant of: R.G. Lye and A.J. Dekker, "Theory of Secondary Emission", Physical Review Vol. 107, pp.977-981 (1957).]:



Now we can proceed to a solution of

$$\frac{\partial \tilde{N}_{ec}(t,E)}{\partial t} = \frac{-1}{\tau(E)} \left[ \tilde{N}_{ec}(t,E) - \int_{E}^{\infty} dE' \frac{\tilde{\delta}_{se}(E')}{E'} \tilde{N}_{ec}(t,E') \right] + \tilde{N}_{pe}(E) \sum_{n=1}^{N_{b}} \delta(t-T_{n})$$
without SE we found
$$\tilde{N}_{ec}(t,E) = \tilde{N}_{pe}(E) \sum_{n=1}^{N_{b}} h(t-T_{n}) \exp\left[-\frac{t-T_{n}}{\tau(E)}\right]$$
with SE the solution is
$$\tilde{N}_{ec}(t,E) = \tilde{N}_{pe}(E) \sum_{n=1}^{N_{b}} h(t-T_{n}) \exp\left[-\frac{t-T_{n}}{\tau_{ec}(E)}\right]$$

wherein

$$\tau_{\rm ec}(E) = \tau(E) + \frac{\exp[-\psi(E)]}{\tilde{N}_{\rm pe}(E)} \int_{E}^{\infty} dE' \exp[\psi(E)] \frac{\tilde{\delta}_{\rm se}(E')}{E'} \tilde{N}_{\rm pe}(E') \tau(E')$$
$$\psi(E) = \int_{0}^{E} dE' \frac{\tilde{\delta}_{\rm se}(E')}{E'}$$

The average density during the train duration is

$$\overline{n}_{ec} \equiv \frac{1}{\left(D_x D_y\right) \left(2\pi\rho\right)} \frac{1}{T N_b} \int_0^\infty dt \int_0^\infty dE \widetilde{N}_{ec}\left(t, E\right)$$
$$= \frac{\overline{N}_{pe}}{\left(D_x D_y\right) \left(2\pi\rho\right)} \frac{1}{T} \int_0^\infty dE f_{pe}\left(E\right) \tau_{ec}\left(E\right)$$

$$\overline{\tau}_{\rm ec} \equiv \int_{0}^{\infty} dE f_{\rm pe}(E) \tau_{\rm ec}(E)$$

For a 5GeV train, the average decay is 10nsec in comparison to about 2.5nsec when no SE were present.
At 2GeV the average decay increases from 10 to 17nsec.
In both cases the average decay is virtually independent of the density.

This result can be attributed to SE being significantly slower than the PE and as a result, it takes them longer time to return to the wall or traverse the gap





Relying on discharge analogy

$$N_{\rm ec}^{\rm (low)} = \frac{\overline{N}_{\rm pe}}{\exp(T/\overline{\tau}_{\rm ec}) - 1}$$

$$N_{\rm ec}^{\rm (high)} \simeq \overline{N}_{\rm pe} + \frac{\overline{N}_{\rm pe}}{\exp(T/\overline{\tau}_{\rm ec}) - 1} = \frac{\overline{N}_{\rm pe}}{1 - \exp(-T/\overline{\tau}_{\rm ec})}$$

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### e-Cloud Static Spatial Distribution



Note that in any case, in the center, the potential well may be approximated by the potential of an harmonic oscillator.

$$\overline{\phi}(\overline{y}) \simeq \overline{\phi}(\overline{y} = 1/2) + \frac{d\overline{\phi}}{d\overline{y}} \int_{\overline{y}=1/2}^{\infty} \delta \overline{y} + \frac{1}{2} \frac{d^2 \overline{\phi}}{d\overline{y}^2} \Big|_{\overline{y}=1/2}^{\infty} \delta \overline{y}^2 \quad \kappa = \frac{d^2 \overline{\phi}}{d\overline{y}^2} \Big|_{\overline{y}=1/2}^{\infty} = -\sum_{n=1}^{\infty} (\pi n)^2 \,\overline{\phi}_n \sin\left(\frac{\pi}{2}n\right)$$

$$Spring coefficient$$
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# e-Cloud Static Spatial Distribution

*Implication on the vertical tune.* Let the oscillatory motion (equilibrium) in the absence of the e-cloud be described by

$$\left[\frac{d^2}{dt^2} + \frac{c^2}{\beta_0^2}\right]\delta y = 0$$

In the presence of the e-cloud

$$\left(\frac{d^2}{dt^2} + \frac{c^2}{\beta_0^2}\right)\delta\overline{y} = \frac{-e}{m\gamma D_y}\left(-\frac{\partial\phi}{\partial y}\right) \simeq \frac{5}{m\gamma D_y^2}\frac{e^2N_{\rm ec}}{4\pi\varepsilon_0 D_z}\delta\overline{y}$$

Implying

$$\frac{1}{\beta^2} = \frac{1}{\beta_0^2} - \frac{5N_{ec}r_e}{\gamma D_y^2 D_z} = \frac{1}{\beta_0^2} - 5n_{ec}\frac{r_e}{\gamma}\frac{D_x}{D_y}$$
$$\frac{\Delta v}{v_0} = \sqrt{1 - 5\frac{n_{ec}r_e\beta_0^2}{\gamma}\frac{D_x}{D_y}} - 1 \approx -\frac{5}{2}\frac{n_{ec}r_e\beta_0^2}{\gamma}\frac{D_x}{D_y} = -Vn_{ec}$$

During the passage of the bunch, temporarily, the distribution changes. Electrons repel the cloud and positron attract it making the potential shallower.

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$$\Gamma^{2} = \left(\frac{\pi^{2} n_{x}^{2}}{D_{x}^{2}} + \frac{\omega_{p}^{2}}{c^{2}}\right) \left(1 - \frac{\omega_{p}^{2}}{\omega^{2}}\right)^{-1}, \ \Gamma a = j\xi$$
$$\kappa_{n_{x}}^{2} \equiv \left(\frac{\pi n_{x} a}{D_{x}}\right)^{2} + \left(\frac{\omega_{p}}{c} a\right)^{2}$$
$$\xi_{0} \equiv \left(\frac{\pi n_{x} a}{D_{x}}\right) \operatorname{coth} \left(\frac{\pi n_{x} a}{D_{x}}\left(\frac{D_{y}}{2a} - 1\right)\right)$$

$$\tan\left(\xi\right) = \frac{\xi_0}{\xi}$$

$$\omega_i^2 = \omega_p^2 \frac{\xi_i^2}{\xi_i^2 + \kappa_{n_x}^2}$$

$$Wakes in the e-Cloud$$

$$Generated Power$$

$$P = \int dx \int dy \int dz J_z E_z$$

$$\overline{P} = \frac{-P}{\eta_0 \left(\frac{q_e c}{D_x}\right)^2} = 2 \left(\frac{a}{D_x}\right) \left(\frac{\omega_p}{c} D_x\right)^2 \sum_{n_x=0}^{\infty} \frac{\sin^2\left(\frac{\pi}{2}n_x\right)}{\sinh^2(\psi_{n_x})} \left(\pi n_x \frac{a}{D_x}\right)^2$$

$$Eigen-frequencies$$

$$\times \sum_i \frac{\xi_i^2 \left[\xi_i^2 + \left(\pi n_x \frac{a}{D_x}\right)^2\right]^{-1}}{\left[\xi_i^2 + \xi_0 \left(1 + \xi_0\right)\right] \left(\xi_i^2 + \kappa_{n_x}^2\right)} \left[\frac{N_b^2}{2} \frac{\operatorname{sinc}^2\left(\frac{1}{2}\omega_i T N_b\right)}{\operatorname{sinc}^2\left(\frac{1}{2}\omega_i T\right)}\right]$$



$$\overline{K} = \frac{f_y(y)D_xc}{\eta_0\left(\frac{q_ec}{D_x}\right)^2 \left[g\left(y,\Delta_y\right)\left(y-\frac{D_y}{2}\right)\right]}$$

$$= 2\left(\frac{\omega_p}{c}D_x\right)^2 \frac{Tc}{a}\sum_{n_x=0}^{\infty} \left(\pi n_x \frac{a}{D_x}\right)^4 \frac{\sin^2\left(\pi n_x \frac{1}{2}\right)}{\sinh^2 \psi_{n_x}}$$

$$\times \sum_i \frac{\xi_i^2 \left[\xi_i^2 + \left(\frac{\pi n_x a}{D_x}\right)^2\right]^{-1}}{\left[\xi_i^2 + \chi_{n_x}\left(1 + \chi_{n_x}\right)\right]\left(\xi_i^2 + \kappa_{n_x}^2\right)} N_b \frac{\operatorname{sinc}(\omega_i T) - \operatorname{sinc}(\omega_i T N_b)}{(\omega_i T)^2 \operatorname{sinc}^2\left(\frac{1}{2}\omega_i T\right)}$$
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Wakes in the e-Cloud



Transverse Kick: "spring coefficient"



The spring coefficient as a function of the cloud density, may be positive in which case the vertical dynamics is <u>unstable</u> or it can become stable for densities that K<0. This behavior is affected by both the cloud density and the cloud thickness









for 1x45 Pattern





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### Experimental Implications

- Radiation generated in the wiggler is the main source of PE.
- *Life-time* of the electrons in the cloud plays crucial role.
- SE extend the life-time of the cloud.
- Relative tune shift proportional to the average cloud density.
- Potential distribution: main source of asymmetry between  $e^{-} \mathcal{A} e^{+}$ .
- "Vacuum Cleaner" may be efficient in particular for high  $N_{ec}$ . Repels ions from the surface.
- Space-charge waves in the cloud generate microwave radiation.
- Radiation and transverse kick peak at resonance  $(T=T_p)$ .
- SC wake treat differently trailing bunches.
- Wake has an impact also in case of an <mark>array</mark> of e-Clouds.



# Photo-Electrons –Preliminary Conclusions

	$E_{\text{th}}[eV]$	$E_k = 2[GeV]$		$E_k = 5[GeV]$	
		$\rho_{\rm eff}\left(E\right) = 0$	$\rho_{\rm eff}\left(E\right) \neq 0$	$\rho_{\rm eff}\left(E\right) = 0$	$\rho_{\rm eff}(E) \neq 0$
$N_{ m pe}$ / $N_e$	0	25.8	20.9	32.9	27.6
	300	0.047(0.18%)	0.047(0.22%)	1.1(3.3%)	1.1(4%)
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	300	13.8(5%)		314 (48%)	
$V_{ m ph}$	0	177[kV]		0.690[MV]	
	300	6.6[kV] (4%)		0.667[MV](97%)	

• Number of PE is one order of magnitude smaller than the photons •Effective voltage associated with PE two orders of magnitude smaller

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