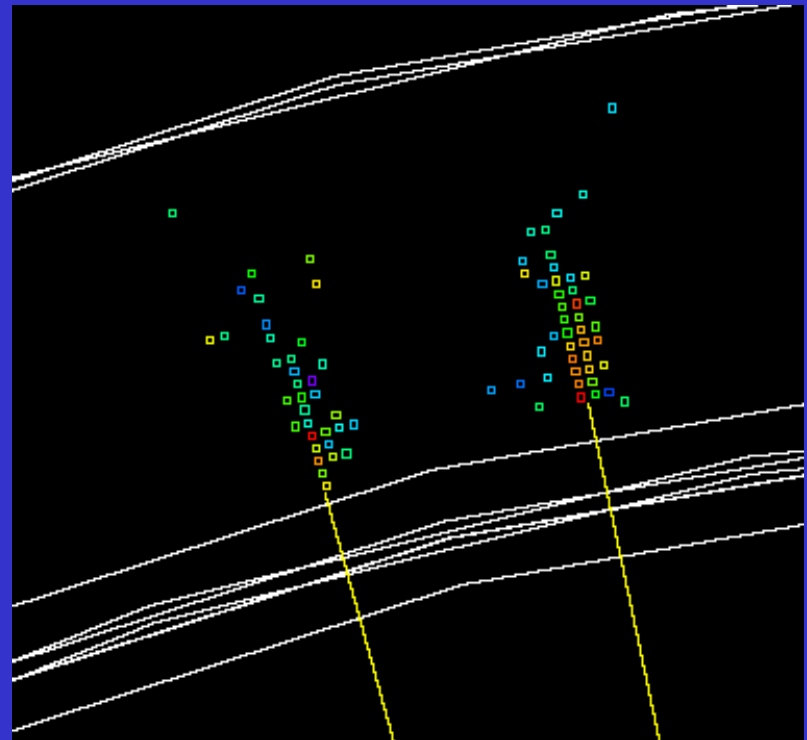


# Using $\pi^0$ mass constraint to improve particle flow ?

Study prompted by looking at event displays like this one of a 5 GeV  $\pi^0$  in sidmay05 detector.

Here photon energies are (3.1, 1.9 GeV), and clearly the photons are very well resolved.

Prompt  $\pi^0$ 's make up most of the EM component of the jet energy.

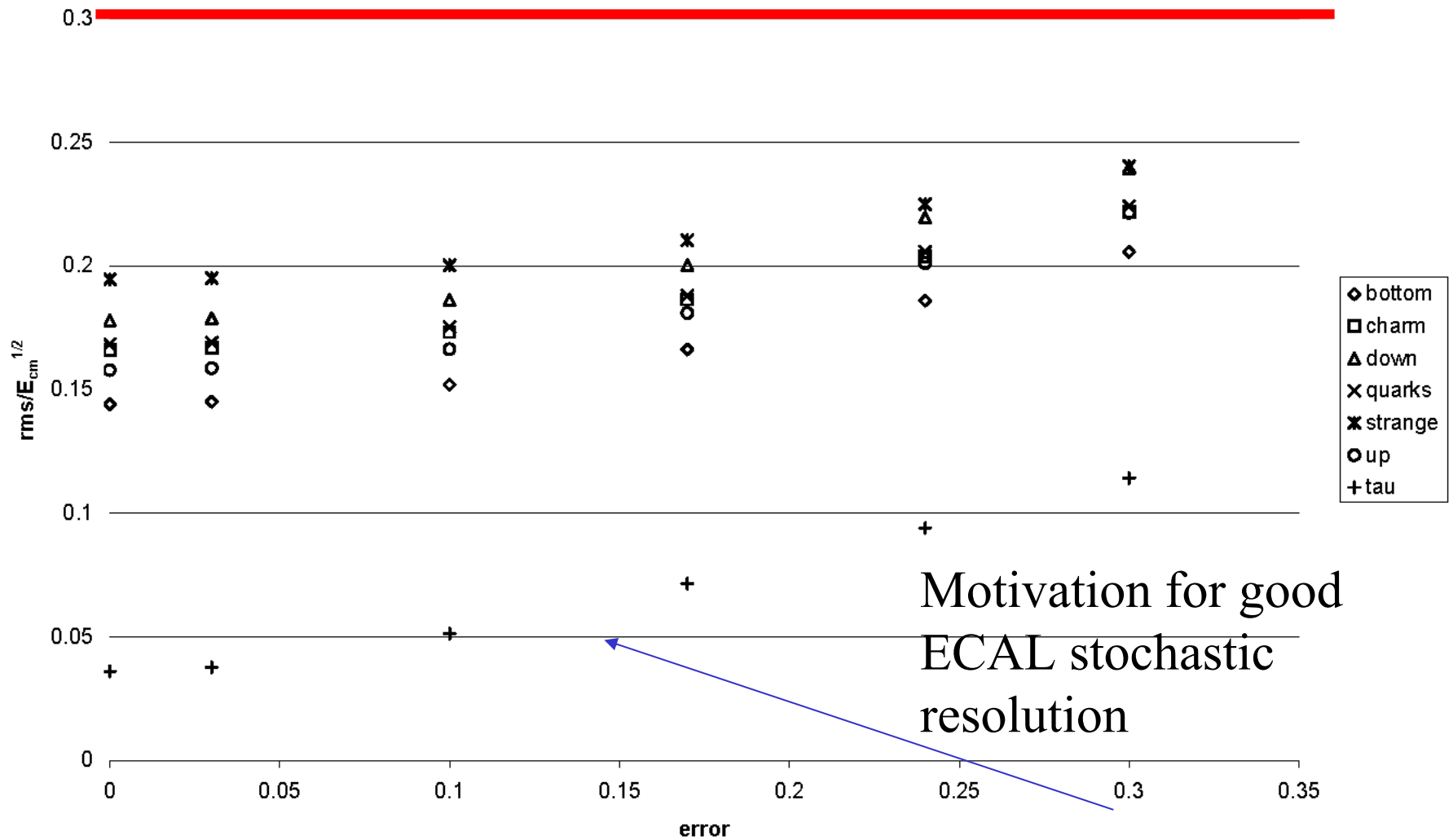


# Investigating $\pi^0$ Kinematic Fits

- Standard technique for  $\pi^0$ 's is to apply the mass constraint to the measured  $\gamma\gamma$  system.
- Setting aside for now the combinatoric assignment problem in jets, I decided to look into the potential improvement in  $\pi^0$  energy measurement.
- In contrast to “normal ECALs”, the Si-W approach promises much better measurement of the  $\gamma\gamma$  opening distance, and hence the opening angle at fixed R. This precise  $\theta_{\gamma\gamma}$  measurement therefore potentially can be used to improve the  $\pi^0$  energy resolution.
- How much ? ( My educated? guess was a factor of  $\sqrt{2}$ ), and how does this affect the detector concepts ?

$C_1$ 

emcal stochastic  
91 GeV



Motivation for good  
ECAL stochastic  
resolution

# Methodology

- Wrote toy MC to generate 5 GeV  $\pi^0$  with usual isotropic CM decay angle ( $dN/d\cos\theta^* = 1$ ).
- Assumed photon energy resolution ( $\sigma_E/E$ ) of  $16\%/\sqrt{E}$ .
- Assumed  $\gamma$ - $\gamma$  opening angle resolution of 2 mrad.
- Solved analytically from first principles, the constrained fit problem under the assumption of a diagonal error matrix in terms of  $(E_1, E_2, 2(1-\cos\theta_{12}))$ , and with a first order expansion.
  - was hoping to get some insight into the analytic dependence on photon resolution assumptions, but problem was somewhat harder than I expected (had to solve a cubic)
  - Note.  $m^2 = 2 E_1 E_2 (1 - \cos\theta_{12})$
- $\pi^0$  kinematics depends a lot on  $\cos\theta^*$ . Useful to define the energy asymmetry,  $a \equiv (E_1 - E_2)/(E_1 + E_2) = \cos\theta^*$ .

# $\pi^0$ mass resolution

- Can show that for  $\sigma_E/E = c_1/\sqrt{E}$  that
$$\Delta m/m = c_1/\sqrt{[(1-a^2) E_{\pi^0}]} \oplus 3.70 \Delta\theta_{12} E_{\pi^0} \sqrt{(1-a^2)}$$

So the mass resolution has 2 terms

- i) depending on the EM energy resolution
- ii) depending on the opening angle resolution

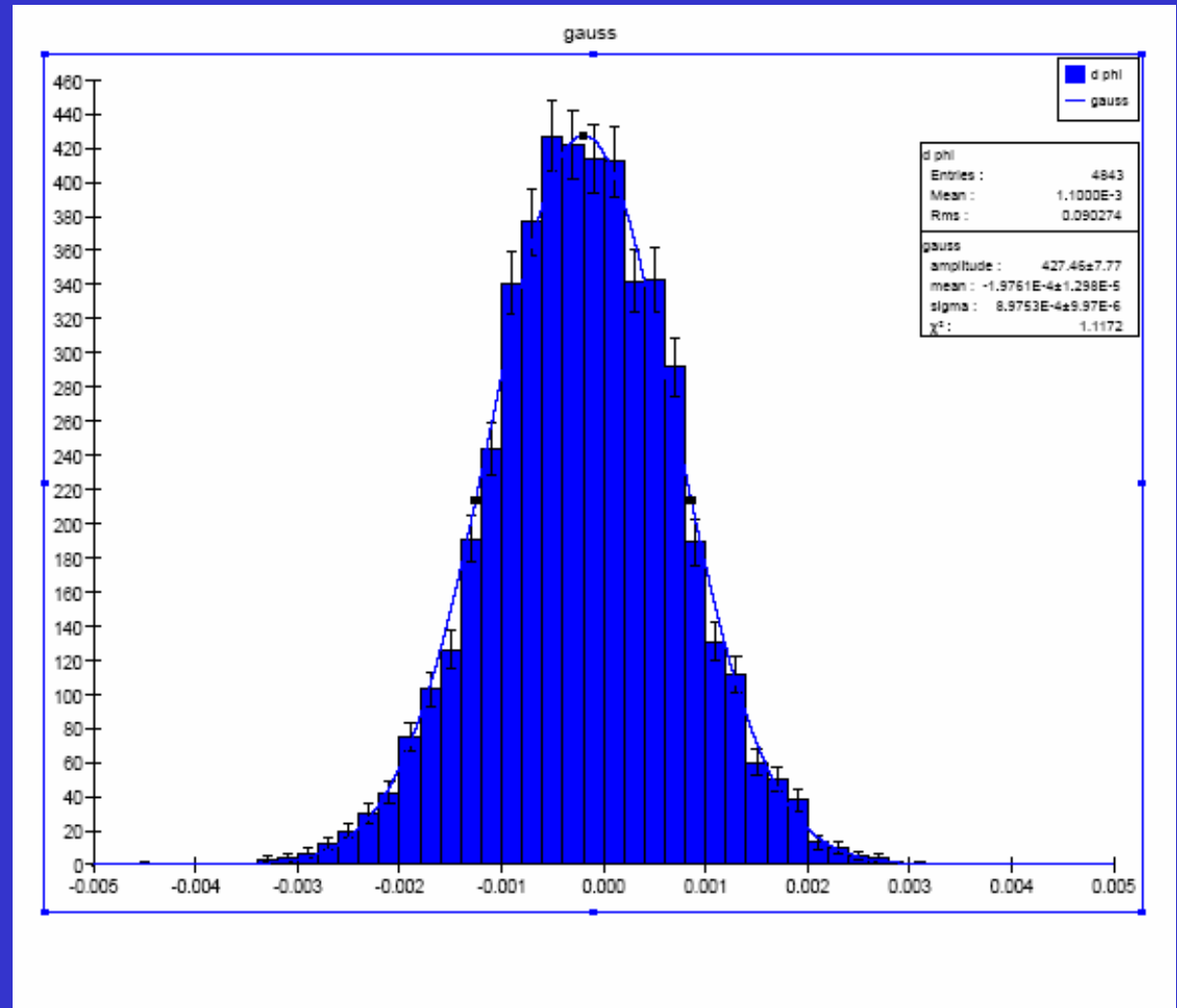
The relative importance of each depends on  $(E_{\pi^0}, a)$

# Angular Resolution Studies

5 GeV photon at  
90°, sidmay05  
detector.

Phi resolution of  
0.9 mrad *just*  
using cluster  
CoG.

=>  $\theta_{12}$  resolution  
of 2 mrad is  
reasonable for  
spatially resolved  
photons.



NB Previous study (see backup slide, shows that a factor of 5 improvement in resolution is possible, (using 1mm pixels !) at fixed R)

# $\pi^0$ mass resolution

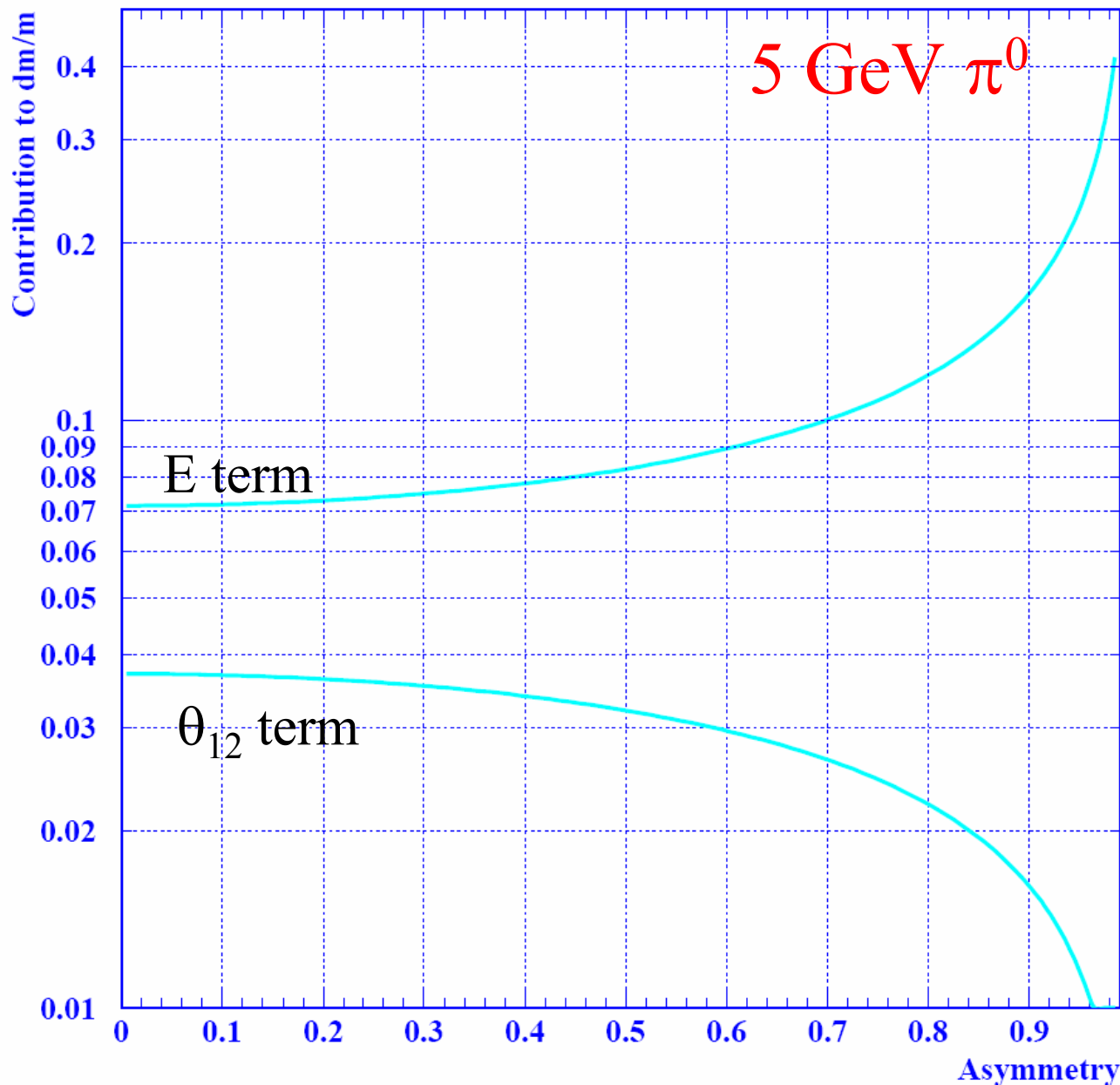
Plots assume:

$$c_1 = 0.16 \text{ (SiD)}$$

$$\Delta\theta_{12} = 2 \text{ mrad}$$

For these detector resolutions, 5 GeV  $\pi^0$  mass resolution dominated by the E term

## pi0 mass resolution contributions



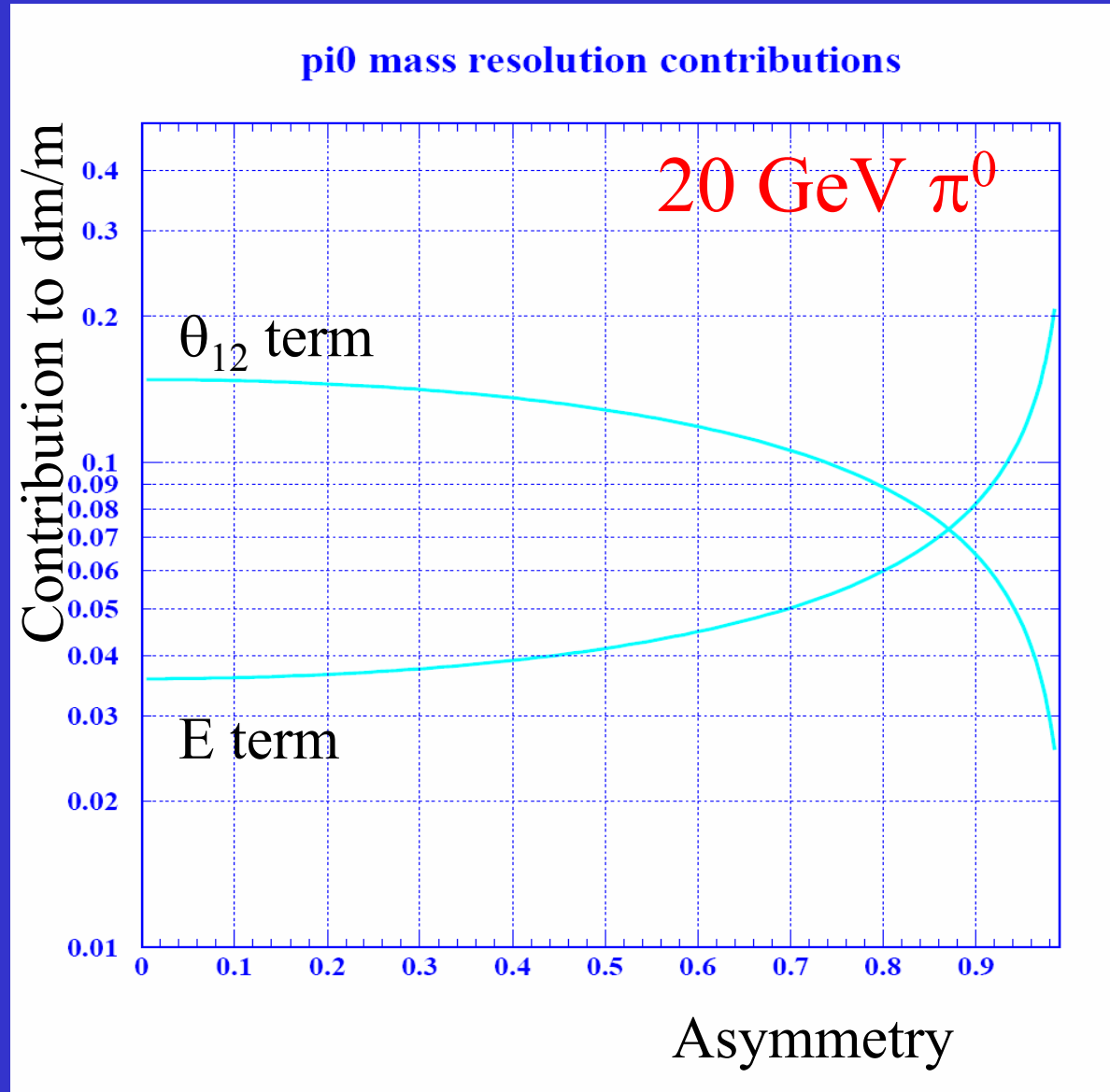
# $\pi^0$ mass resolution

Plots assume:

$$c_1 = 0.16 \text{ (SiD)}$$

$$\Delta\theta_{12} = 2 \text{ mrad}$$

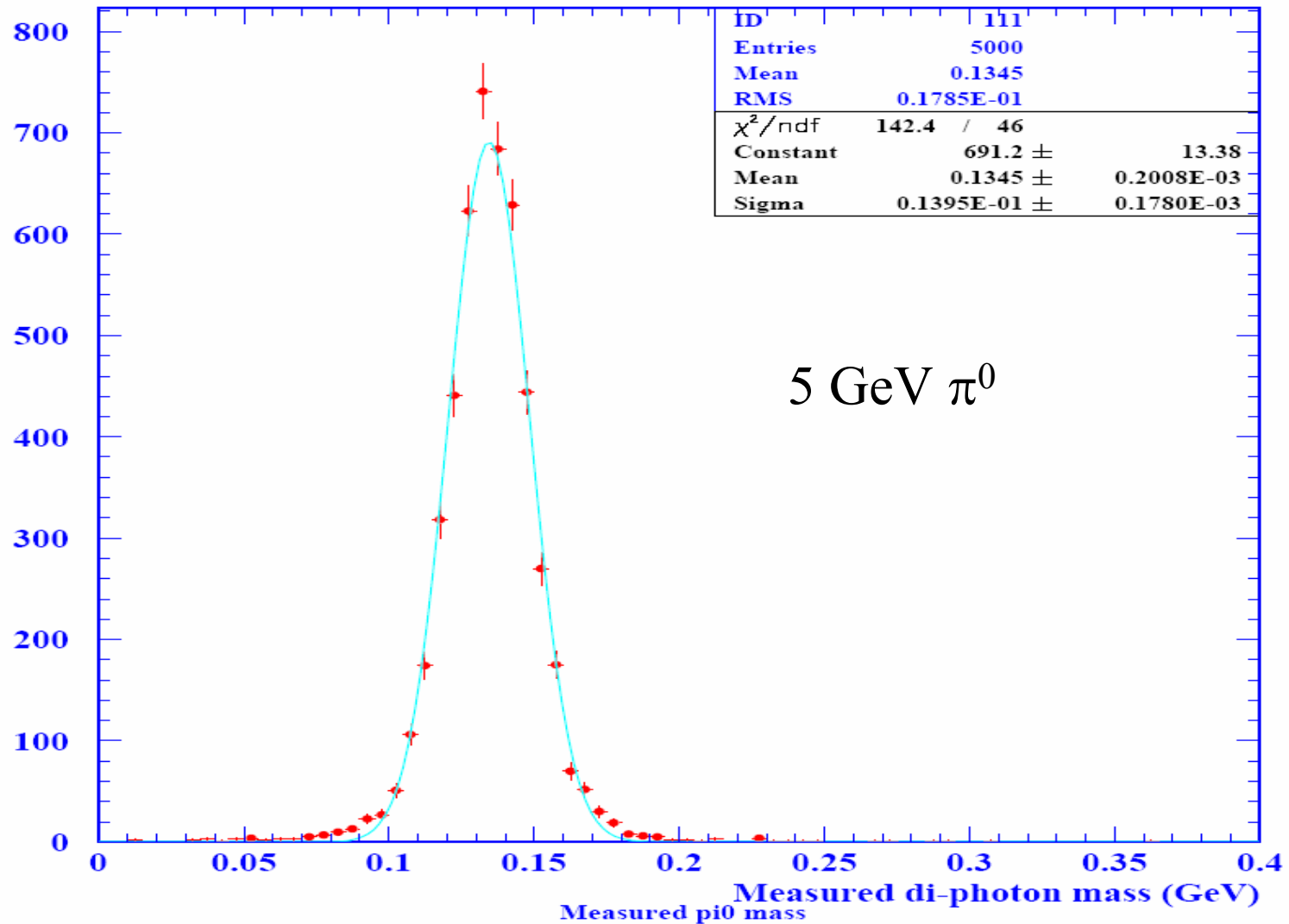
For these detector resolutions, 20 GeV  $\pi^0$  mass resolution dominated by the  $\theta_{12}$  term ( $\Rightarrow$  KF less helpful)





# $\pi^0$ mass

## pi0 kinematic fit



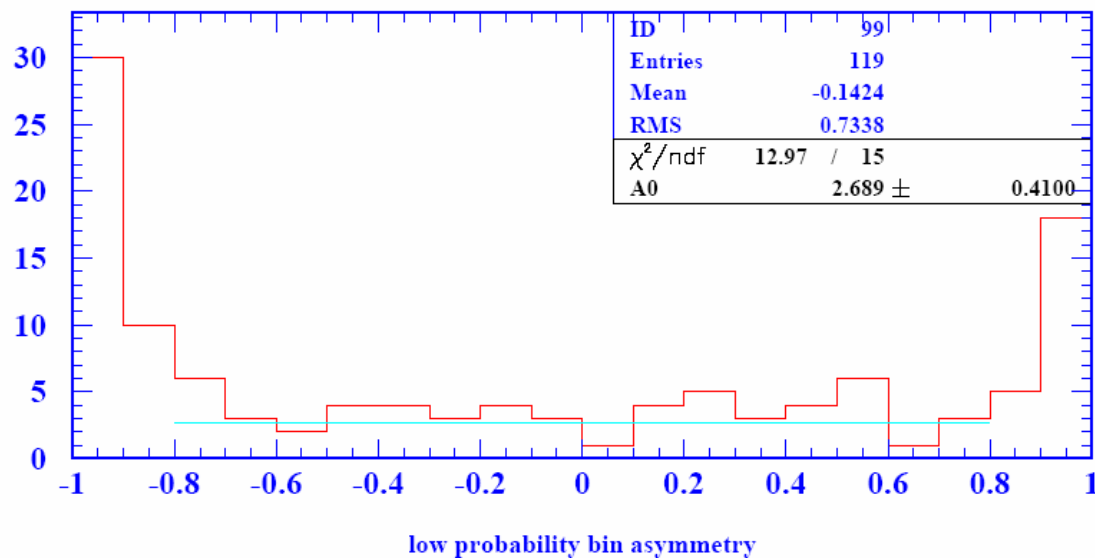
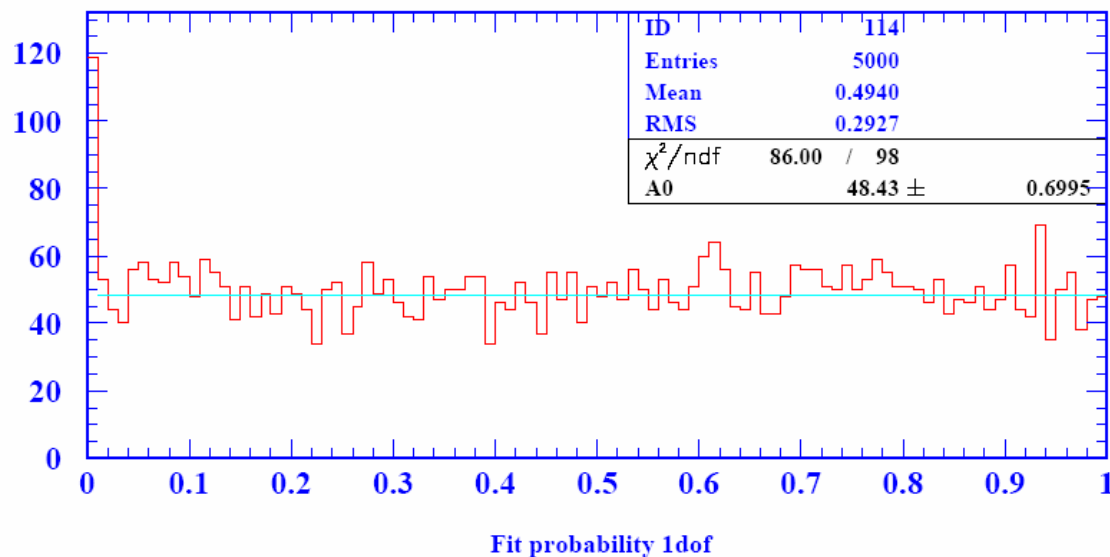
# Fit quality

Probability distribution flat (as expected).

$$a = (E_1 - E_2) / (E_1 + E_2)$$

Spike at low probability corresponds to asymmetric decays ( $|a| \approx 1$ ). I think I need to iterate using the fitted values for the error estimation .....

pi0 kinematic fit

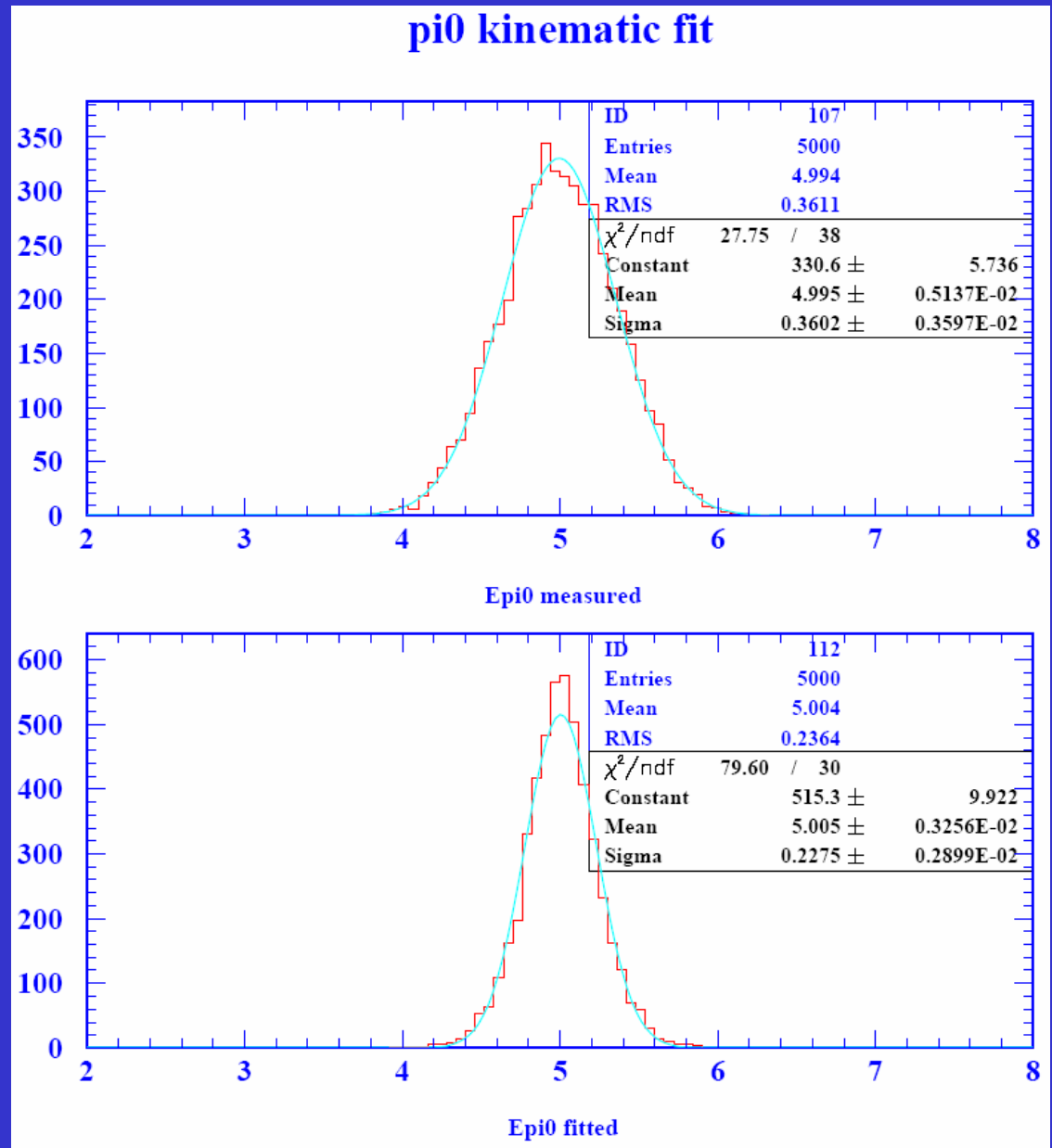


# $\pi^0$ energy

Measured

Fitted (improves  
from 0.36 GeV to  
0.23 GeV)

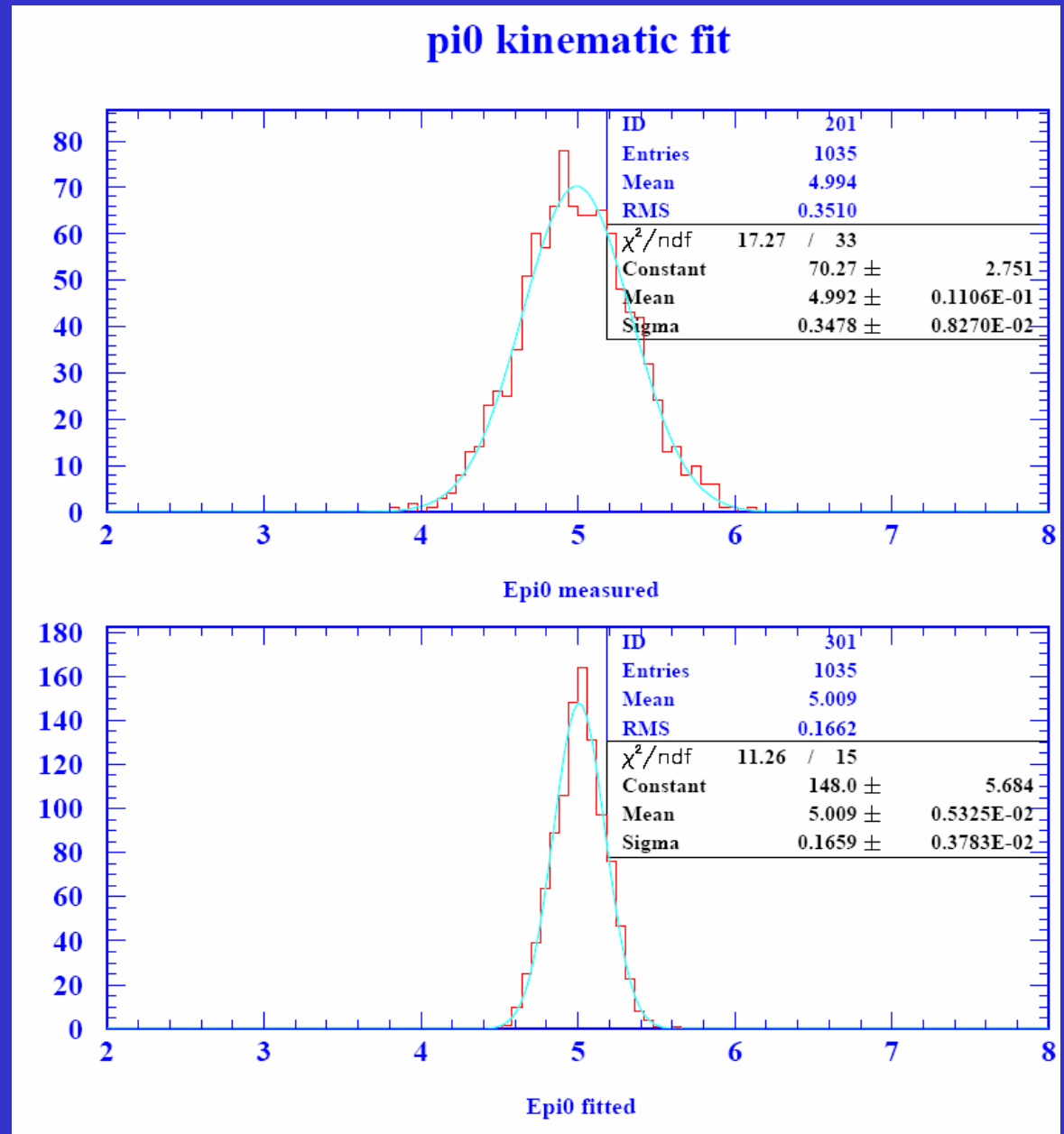
(factor of 0.64 !!)



$\pi^0$  energy  
for  $|a| < 0.2$

Improvement  
most dramatic :

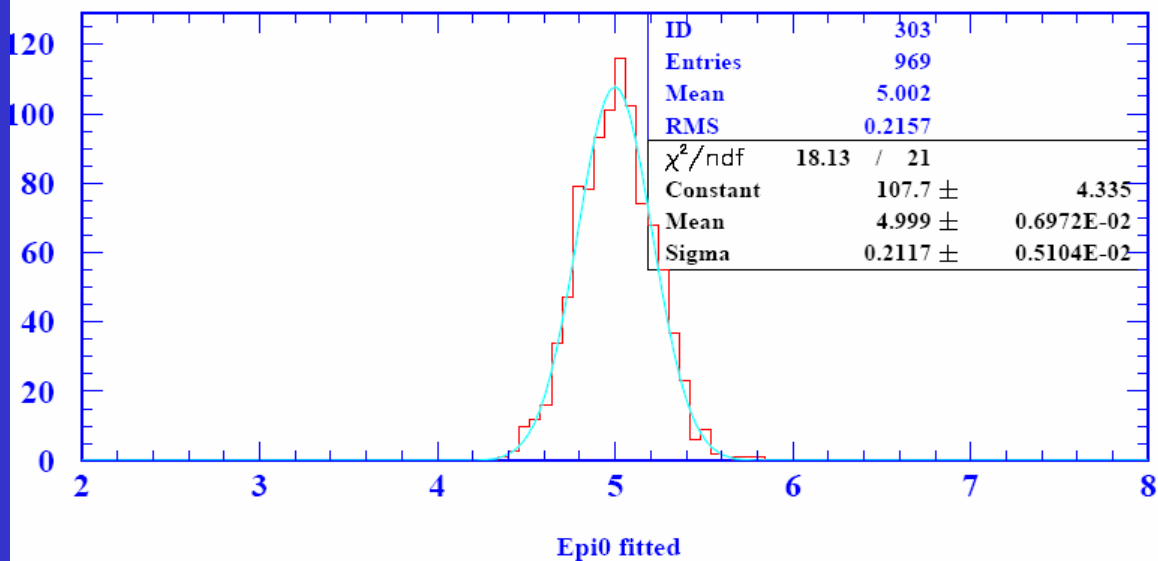
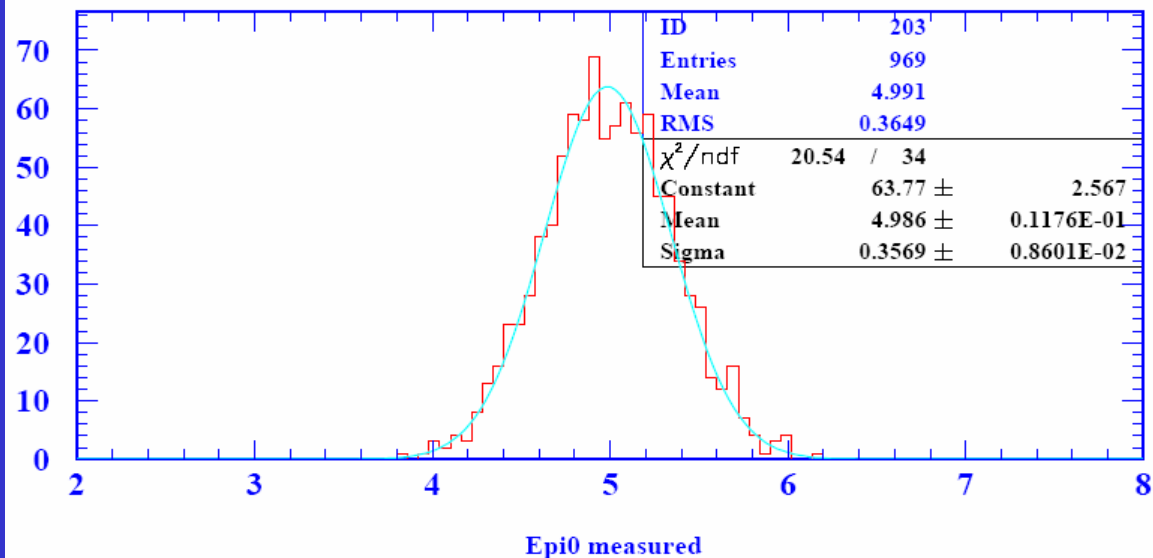
0.35  $\rightarrow$  0.17



$\pi^0$  energy for  
 $0.4 < |a| < 0.6$

Improvement  
from 0.36 to  
0.21

pi0 kinematic fit

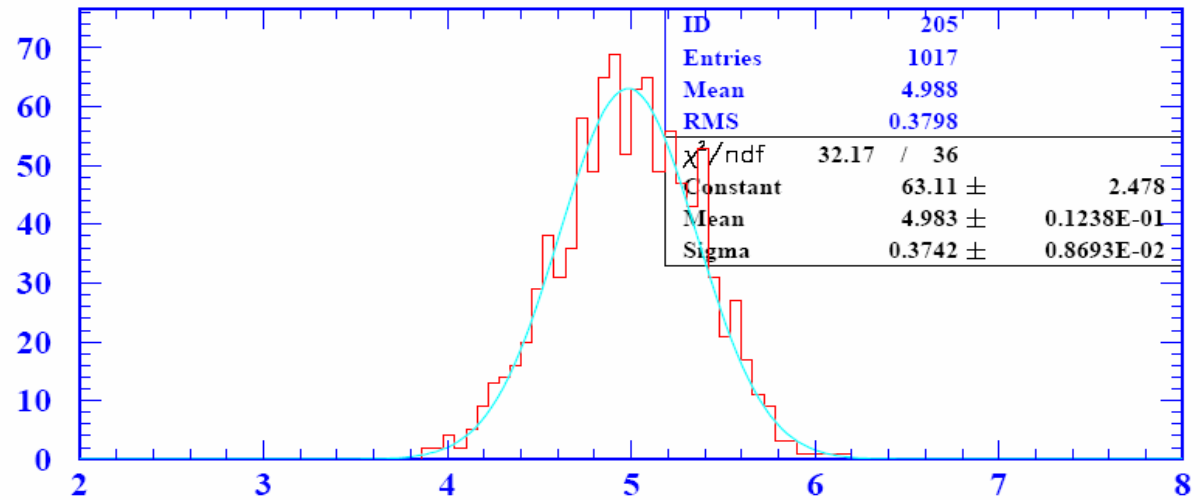


# $\pi^0$ energy for $|a| > 0.8$

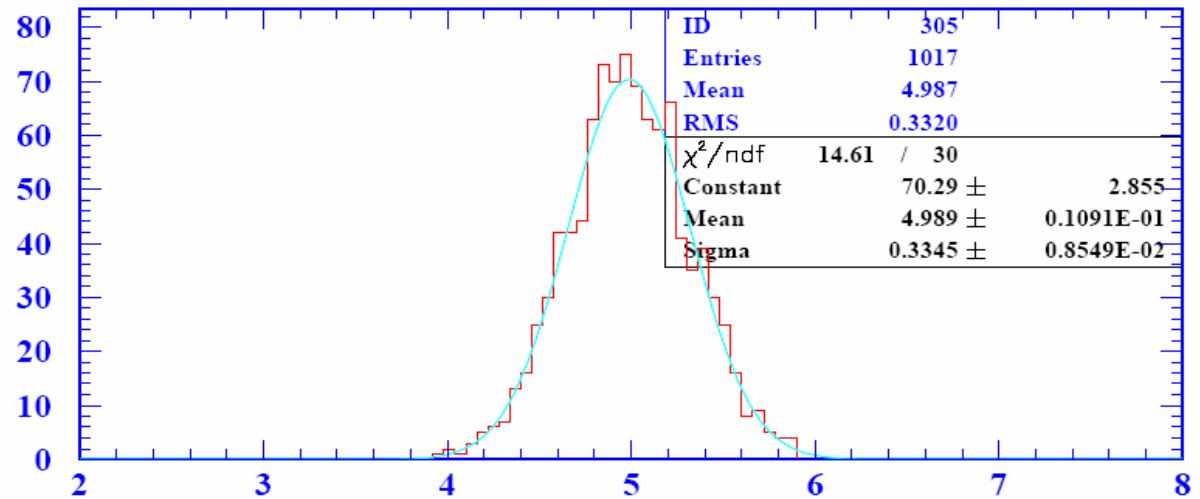
Improvement  
not so great.  
(as expected)

0.37  $\rightarrow$  0.33

### $\pi^0$ kinematic fit



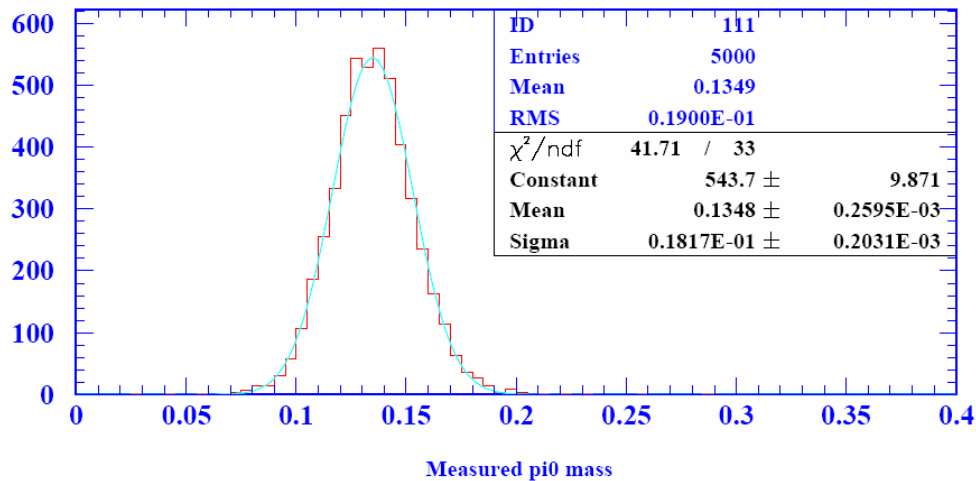
### $E_{\pi^0}$ measured



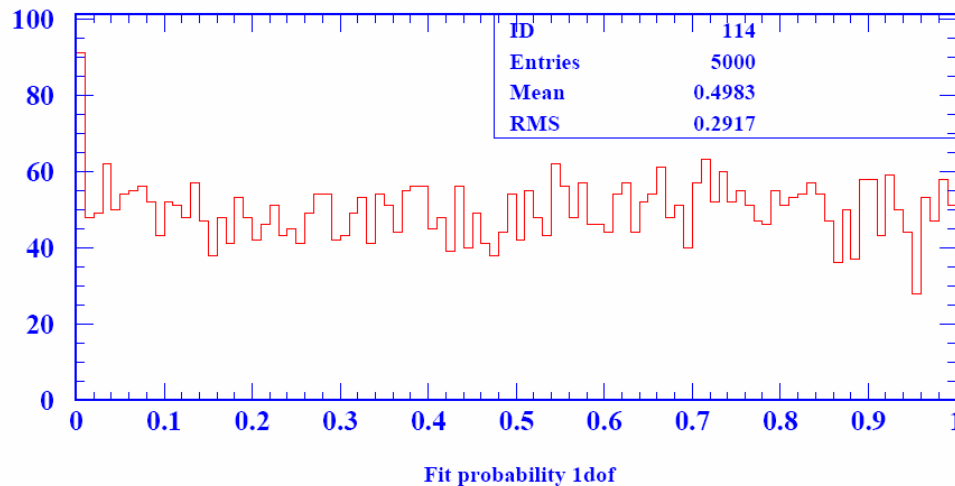
### $E_{\pi^0}$ fitted

# 20 GeV $\pi^0$ , same resolution assumptions

20 GeV pi0 study



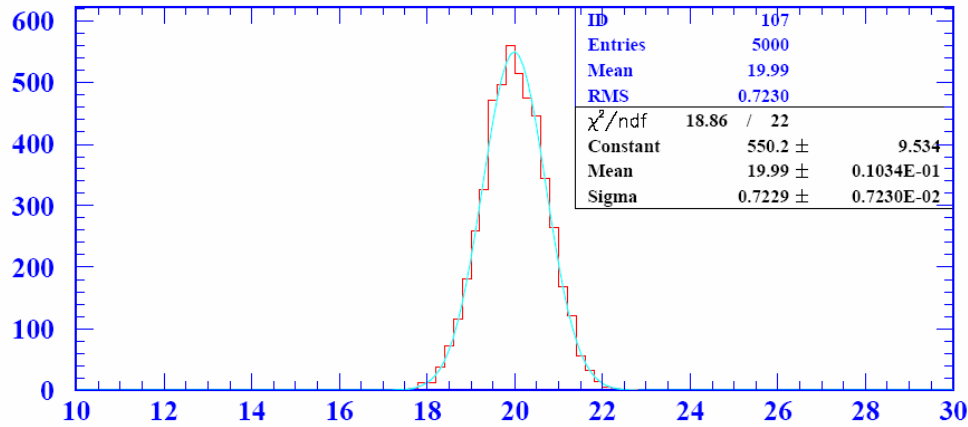
Mass resolution degrades as expected.



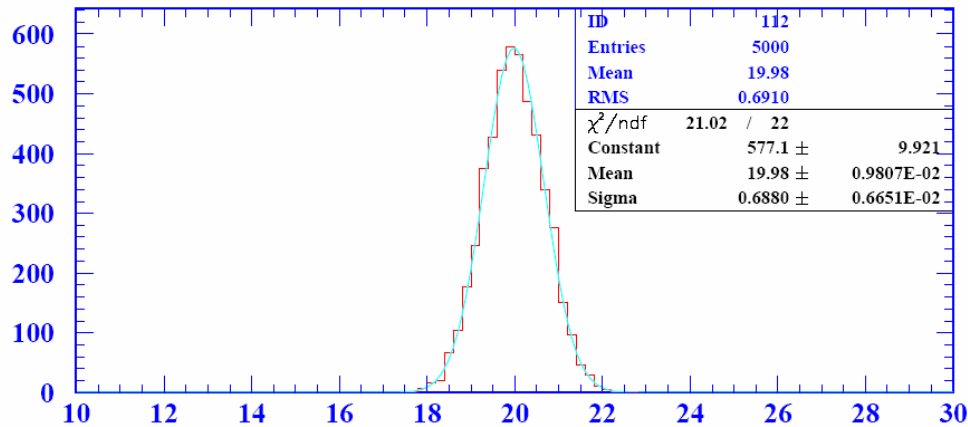
Constrained fit still works OK.

# 20 GeV $\pi^0$ , same resolution assumptions

## 20 GeV pi0 study



Epi0 measured



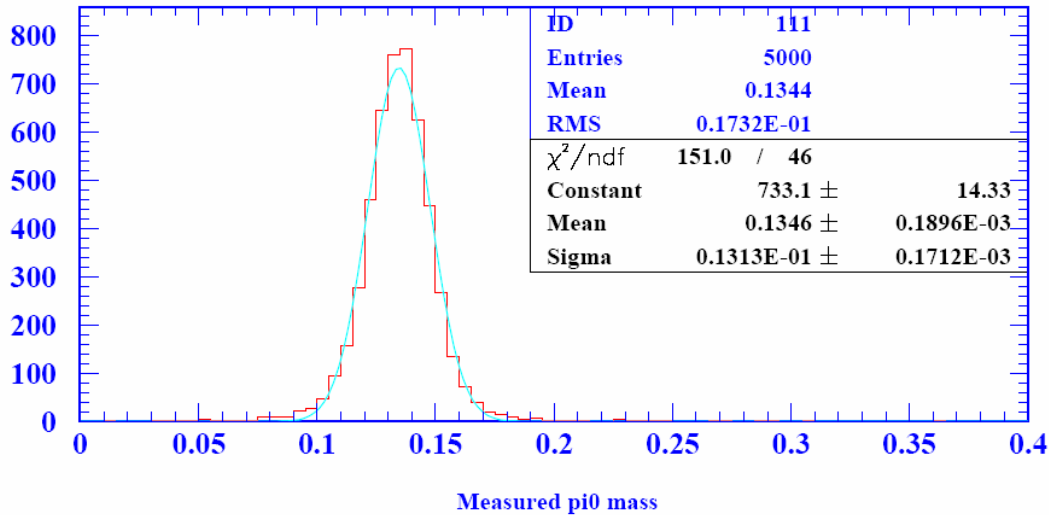
Epi0 fitted

Constrained fit  
 $\Rightarrow$  No significant improvement.  
(as expected)

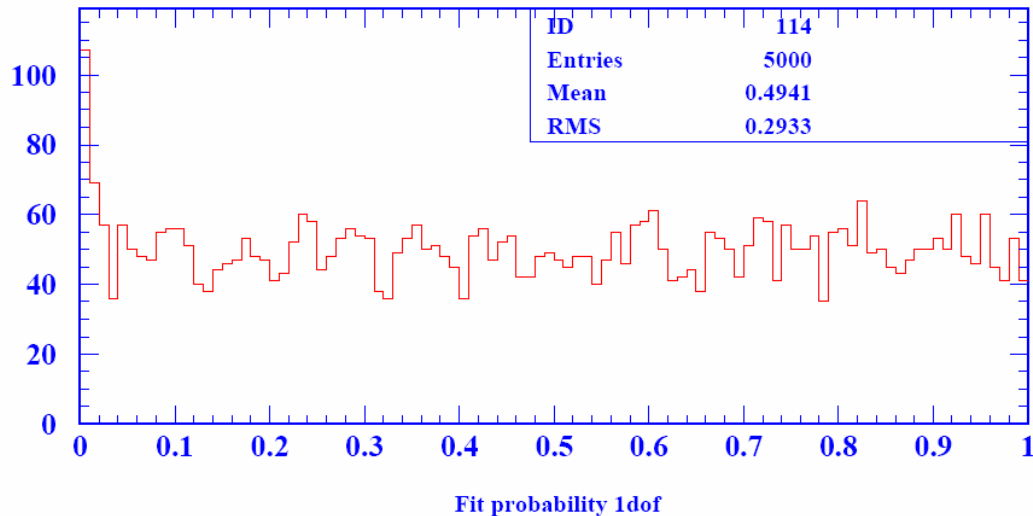


# 5 GeV $\pi^0$ , 4 times better $\theta_{12}$ resolution

5 GeV pi0, 0.5 mrad opening angle resolution



Not much change  
in mass resolution  
(dominated by E-  
term)



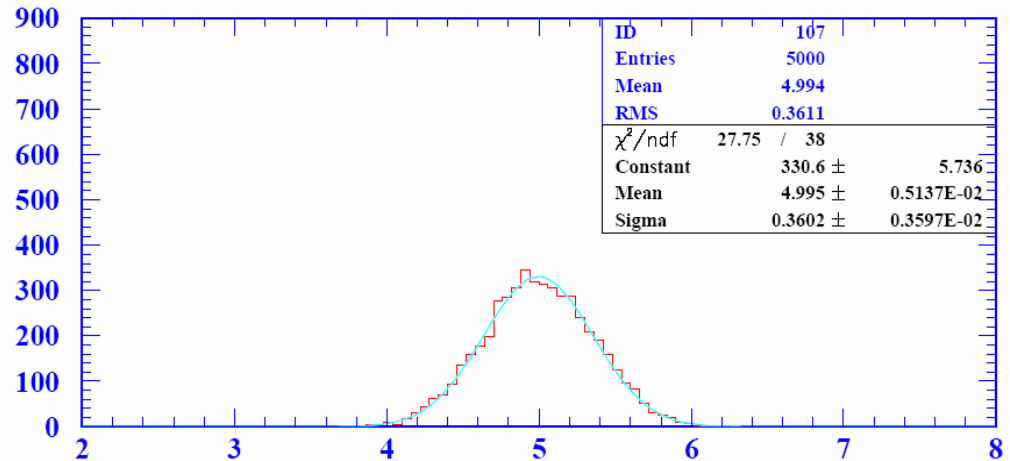
Fit still works.

# $\pi^0$ energy resolution improvement

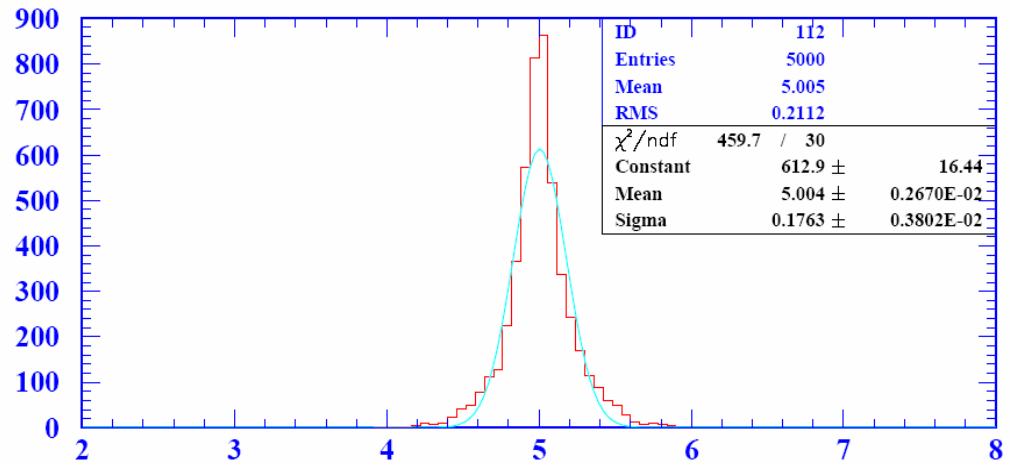
Dramatic !

Factor of 2 for  
ALL asymmetries.

5 GeV  $\pi^0$ , 0.5 mrad opening angle resolution



Epi0 measured



Epi0 fitted

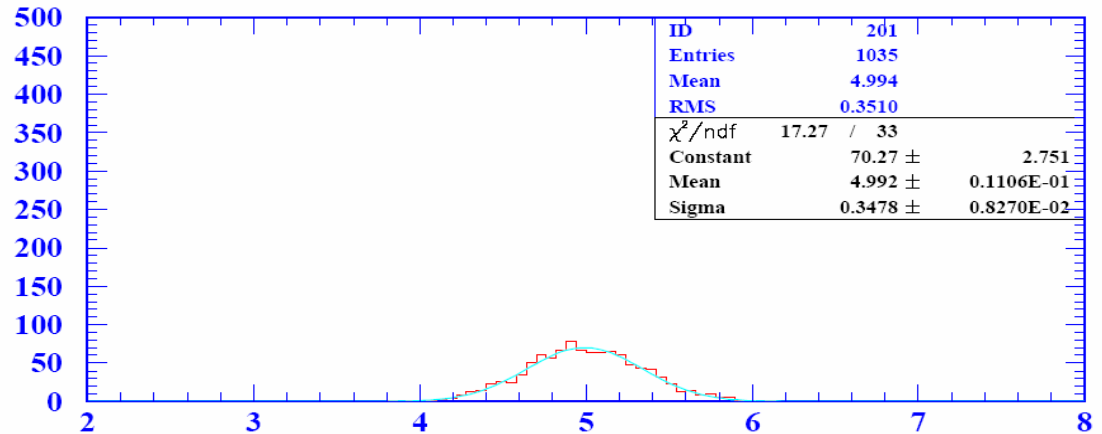
# $\pi^0$ energy resolution improvement

$$|a| < 0.2$$

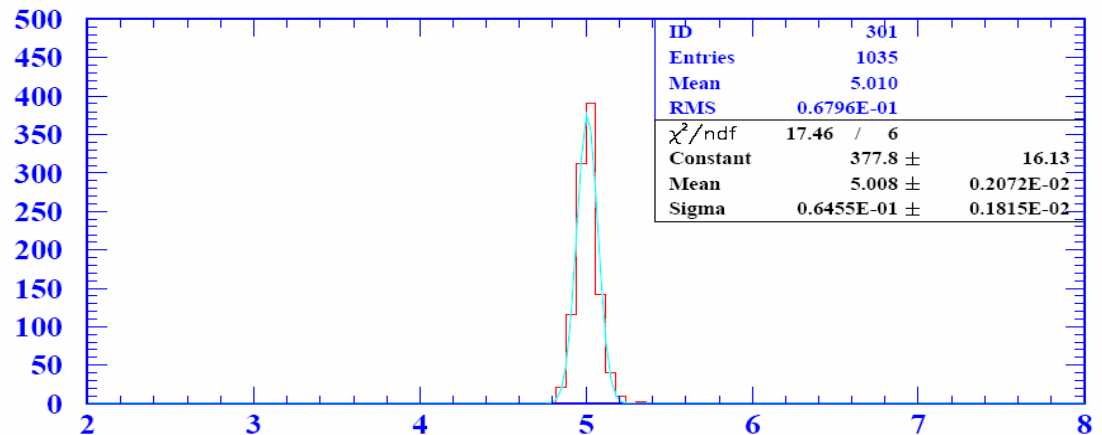
Improves by a factor of 0.35/0.065.

i.e. a factor of 5 !

5 GeV pi0, 0.5 mrad opening angle resolution



Epi0 measured



Epi0 fitted

# Conclusions

- $\pi^0$  constrained fit has a lot of potential to improve the  $\pi^0$  energy resolution.
- Will investigate in more detail actual  $\gamma$ - $\gamma$  separation capabilities.
  - Puts a high premium on angular resolution if this is as useful as it looks.
- Looks worthwhile to also look into the assignment problem.
- May have some mileage for reconstructing the  $\pi^0$ 's in hadronic interactions.

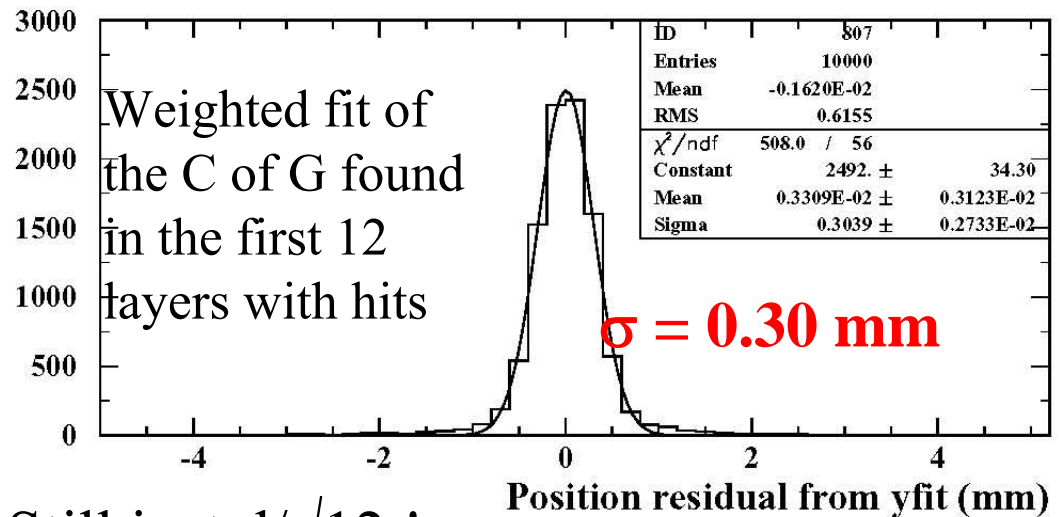
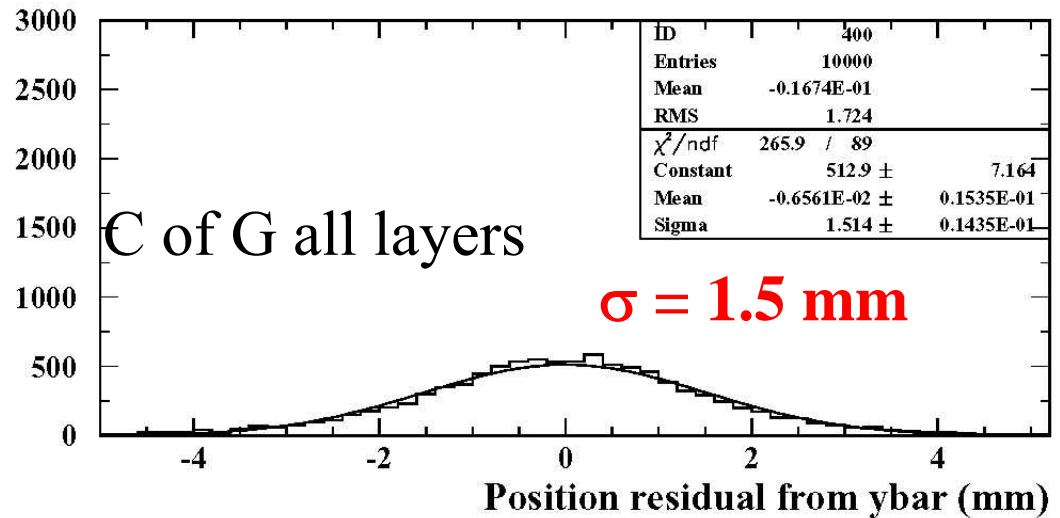
# Backups

# Position resolution from simple fit

Neglect layer 0 (albedo)

Using the first 12 layers with hits with  $E > 180$  keV, combine the measured C of G from each layer using a least-squares fit (errors varying from 0.32mm to 4.4mm). Iteratively drop up to 5 layers in the “track fit”.

*Position resolution does indeed improve by a factor of 5 in a realistic 100% efficient algorithm!*



Still just  $d/\sqrt{12}$  !

hcal stochastic  
91 GeV

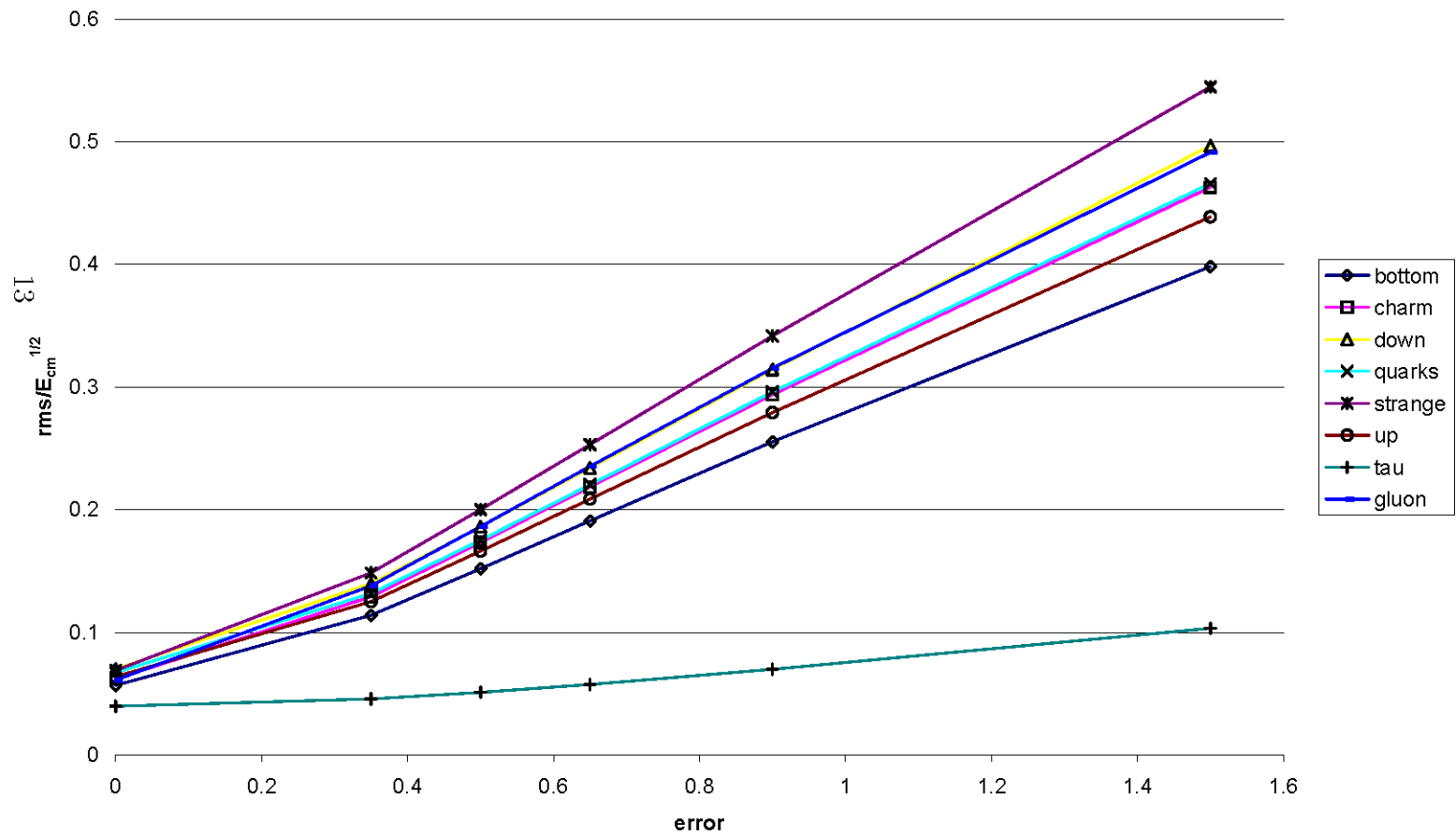


Figure 4