
First Results with the Prototype Detectors of the Si/W ECAL

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- Physics Design Requirements
- Detector Concept
- Silicon Detectors - Capacitance and Trace Resistance
- Implications of Accelerator Technology Choice
- MIPS, sources and laser

Si-W work – personnel and responsibilities

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Simulation

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* This work includes contributions from Oregon students Tyler Neely and Eric Fitzgerald.

ECAL Design Requirements

- Optimal contribution to the reconstruction of multijet events:
 - Excellent separation of γ 's from charged particles
Efficiency > 95% for energy flow
 - Excellent linkage of ECAL with tracker (important for SiD)
 - Good linkage of ECAL with HCAL
 - Good reconstruction of π^\pm , detection of neutral hadrons
 - Reasonable EM energy resolution ($< 15\%/\sqrt{E}$)

Physics case: jet reconstruction important for many physics processes.

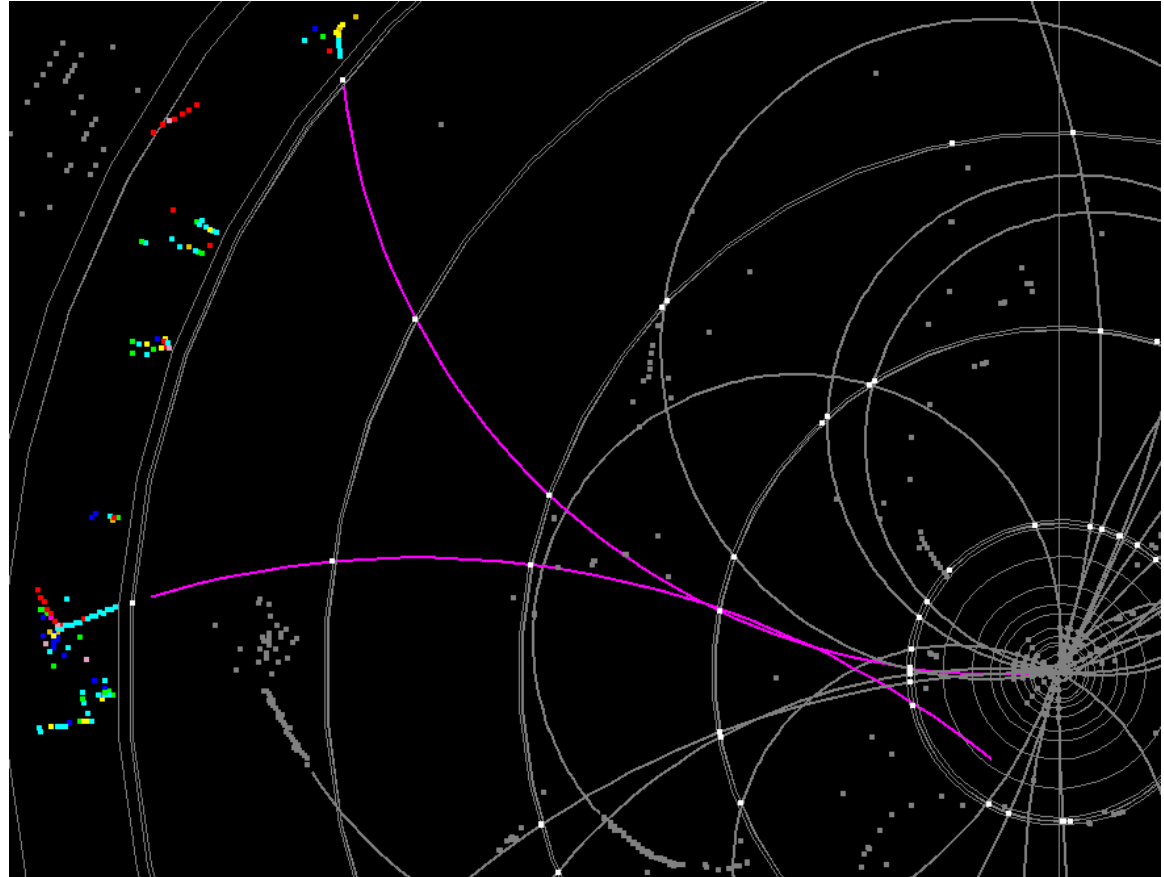
- Longitudinal Sampling, 30 layers needed for EM energy resolution

$$\frac{\sigma_E}{E} \sim 20\% \sqrt{\frac{X}{E}}$$

X is the sampling in radiation length.

- Useful for K^0 tracking, etc.

- **Can tolerate small, random inefficiency**



See talks by Eckhard von Toerne

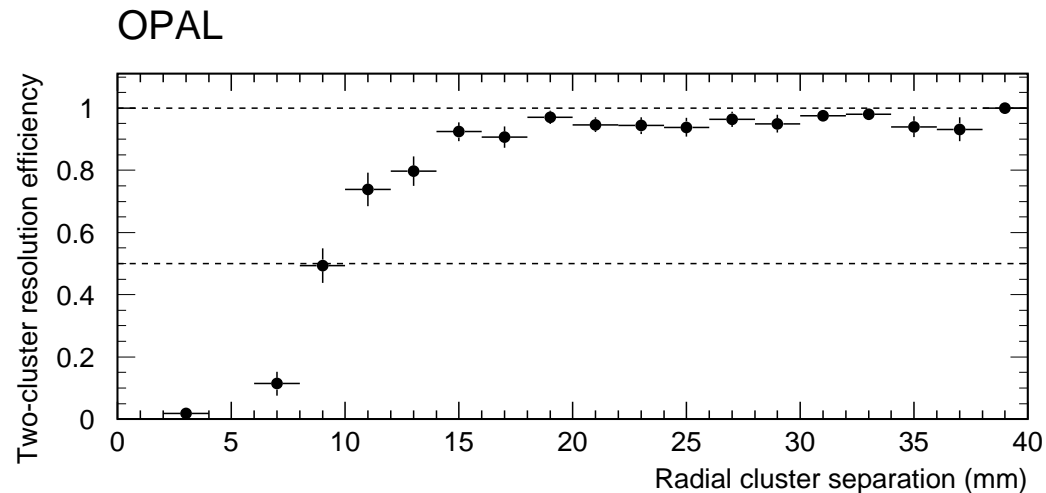
Importance of Granularity

- Figure of merit for energy reconstruction is

$$f_E \simeq \frac{\max(R_M, 4d)}{R_{cal}}$$

where R_M is the Molière radius, d is the detector pad size and R_{cal} is the inner radius of the calorimeter (factor of 4 somewhat arbitrary)

Example (OPAL SiW luminosity monitor, $1X_0$ radiator, 3mm gap)



$$d = 2.5\text{mm} , R_M \sim 17\text{mm}$$

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- The costs of the calorimeters, coil, and muon system have

$$\text{cost} \propto R_{cal}^n$$

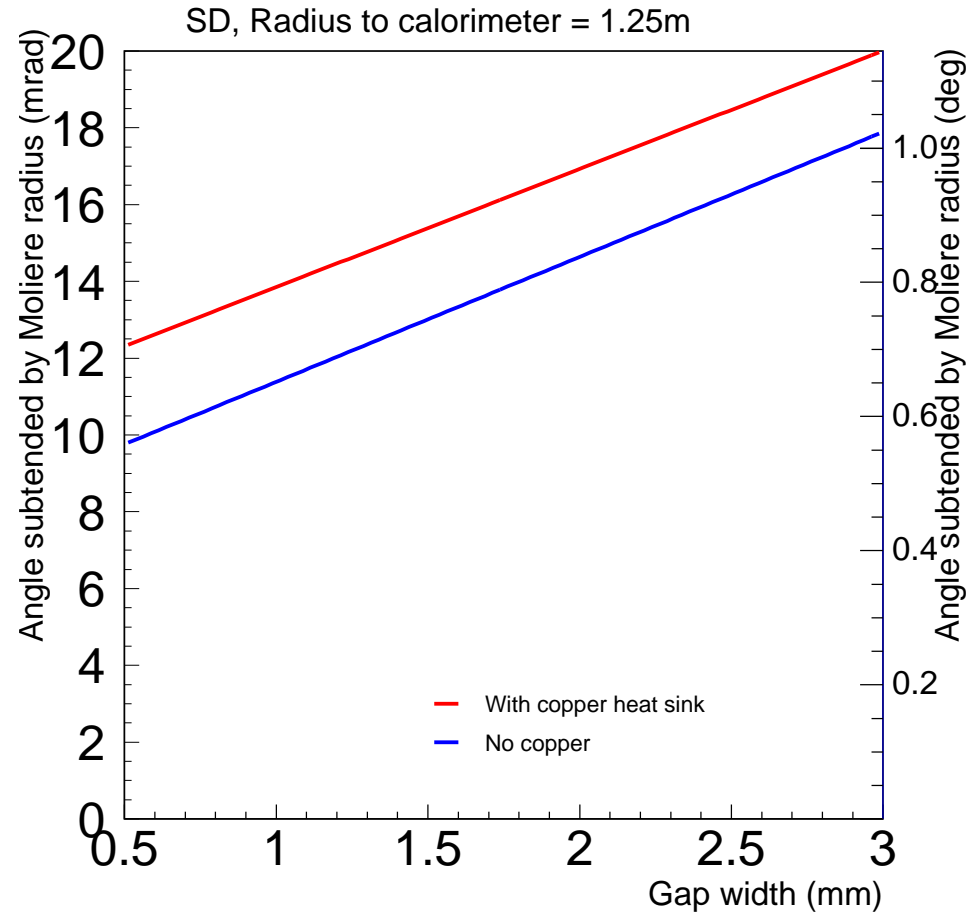
where n is $\sim 2 - 3$.

- Thus a 10% increase in the Molière radius of the calorimeter leads to a $> 20\%$ increase in cost of the detector for constant f_e .
- Conclusion: try and make the calorimeter as dense as possible

Critical parameter: gap between tungsten layers.

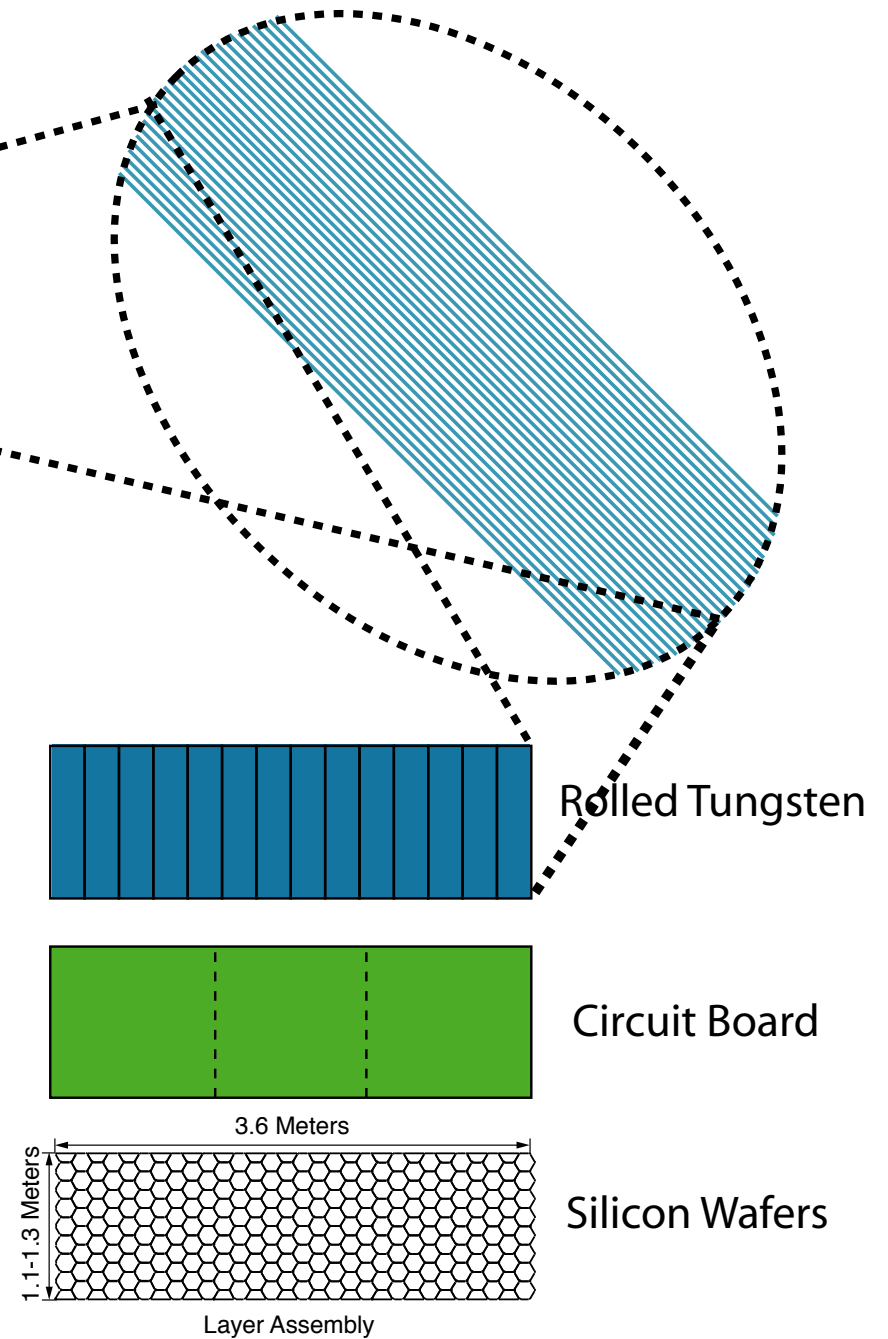
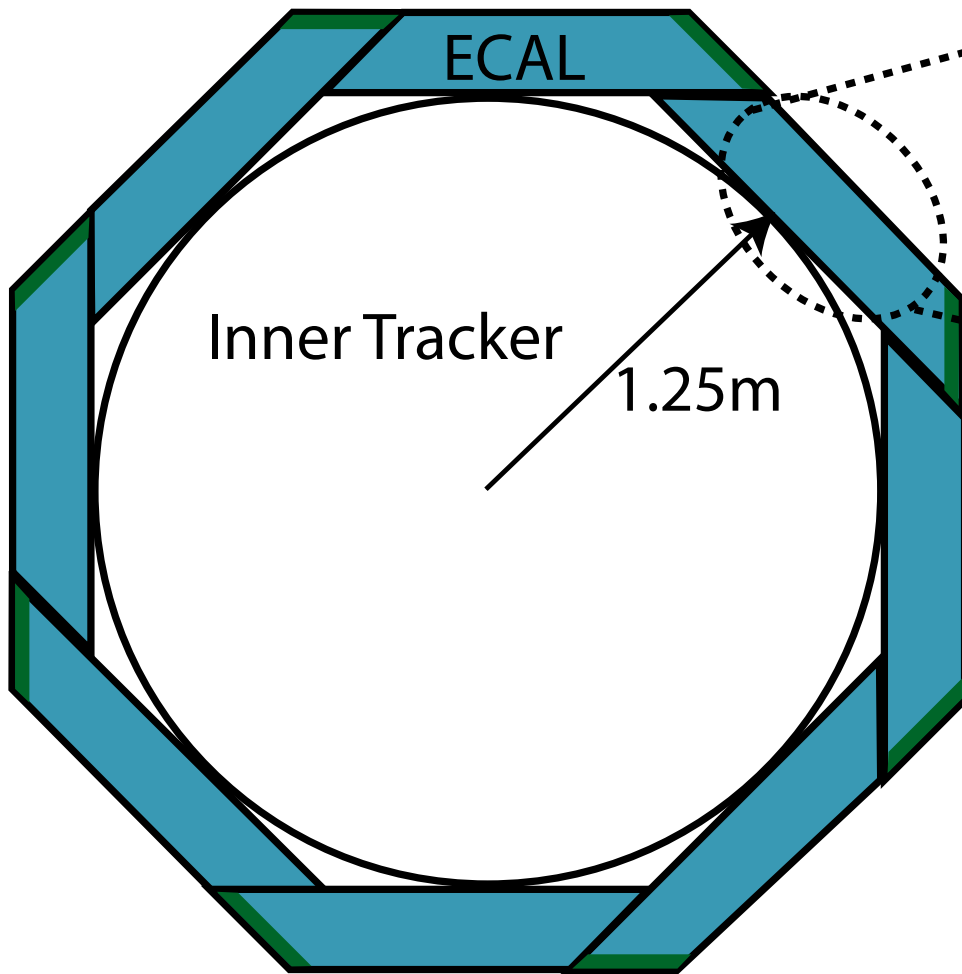
Config.	Radiation length	Molière Radius
100% W	3.5mm	9mm
92.5% W	3.9mm	10mm
+1mm gap	5.5mm	14mm
+1mmCu	6.4mm	17mm

Assumes 2.5mm thick tungsten absorber plates



Calice 3mm gap with 1.7m TESLA radius gives $\frac{R_M}{R_{Cal}} = 13\text{mrad}$

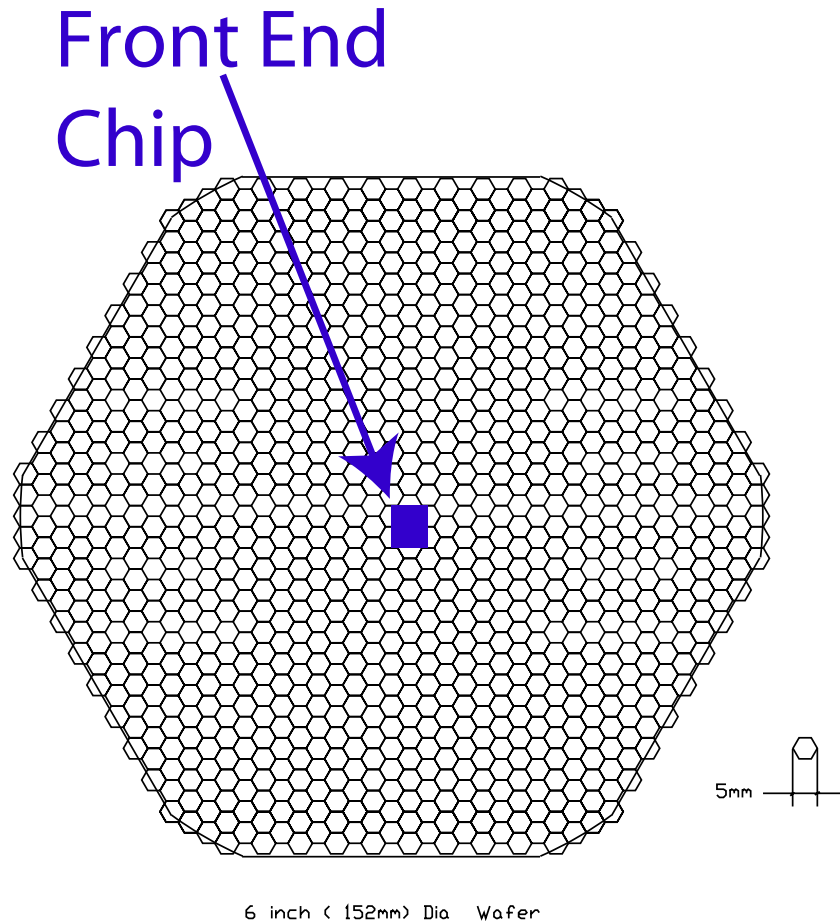
Si-W Calorimeter Concept



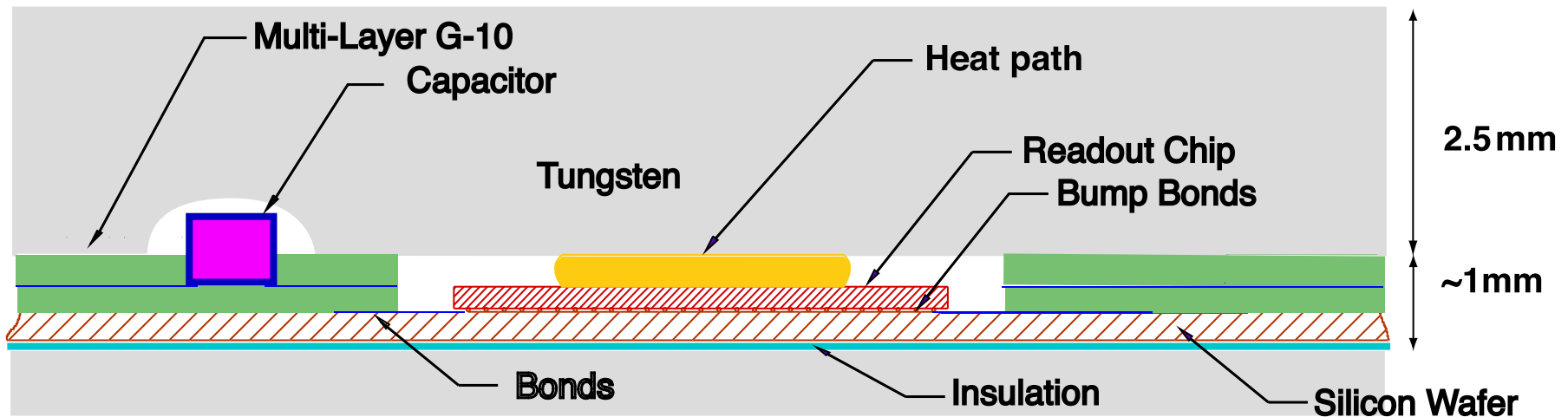
Transverse Segmentation $\sim 5\text{mm}$
30 Longitudinal Samples
Energy Resolution $\sim 15\%/E^{1/2}$

Silicon Concept

- Readout each wafer with a single chip
- Bump bond chip to wafer
- To first order cost independent of pixels /wafer
- Hexagonal shape makes optimal use of Si wafer
- Channel count limited by power consumption and area of front end chip
- May want different pad layout in forward region



Critical parameter: minimum space between tungsten layers.



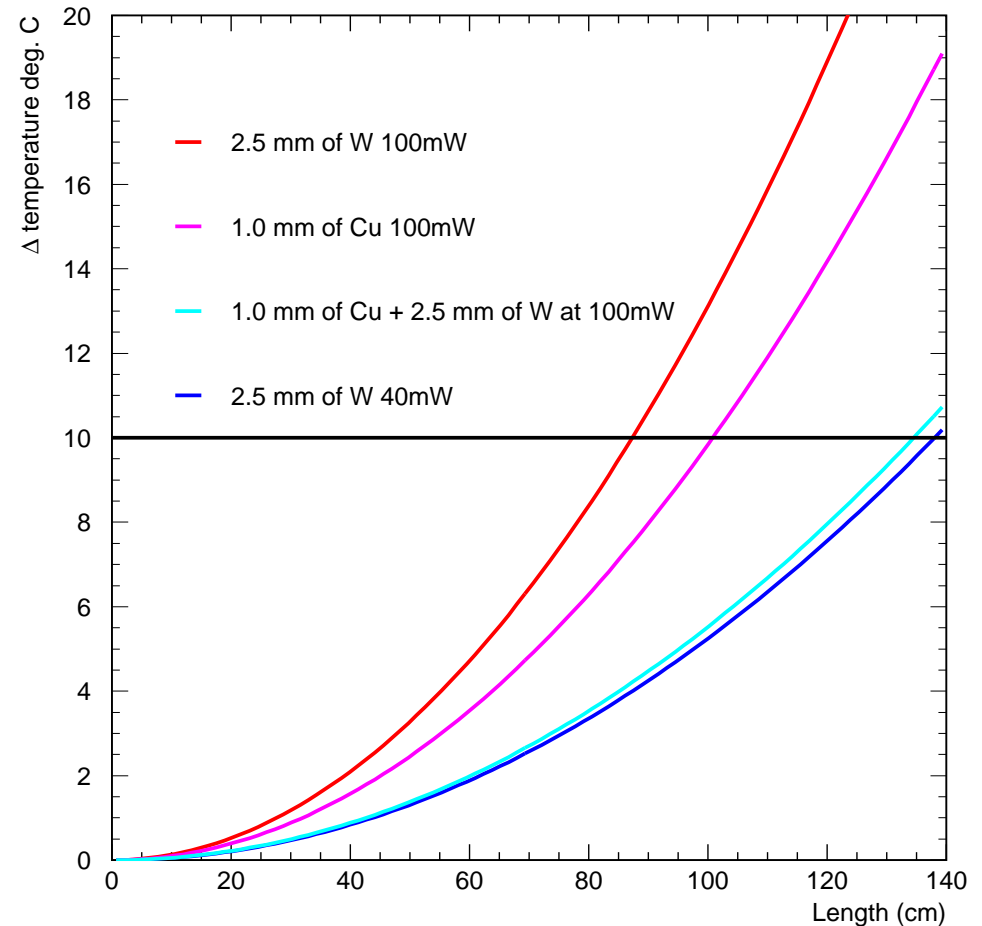
Evolving capacitor packaging may eliminate need for dimples.

Can we get the heat out?

Back of the envelope calculation of change in temperature:

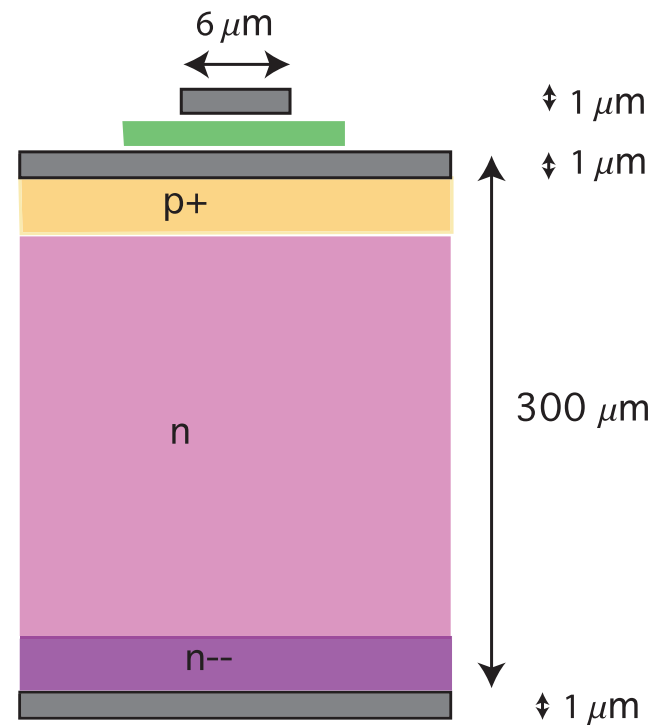
- Thermal Conductivity of W alloy $120\text{W}/(\text{K}\cdot\text{m})$
- Thermal Conductivity of Cu $400\text{W}/(\text{K}\cdot\text{m})$

Need to reduce heat to below $100\text{mW}/\text{wafer}$.

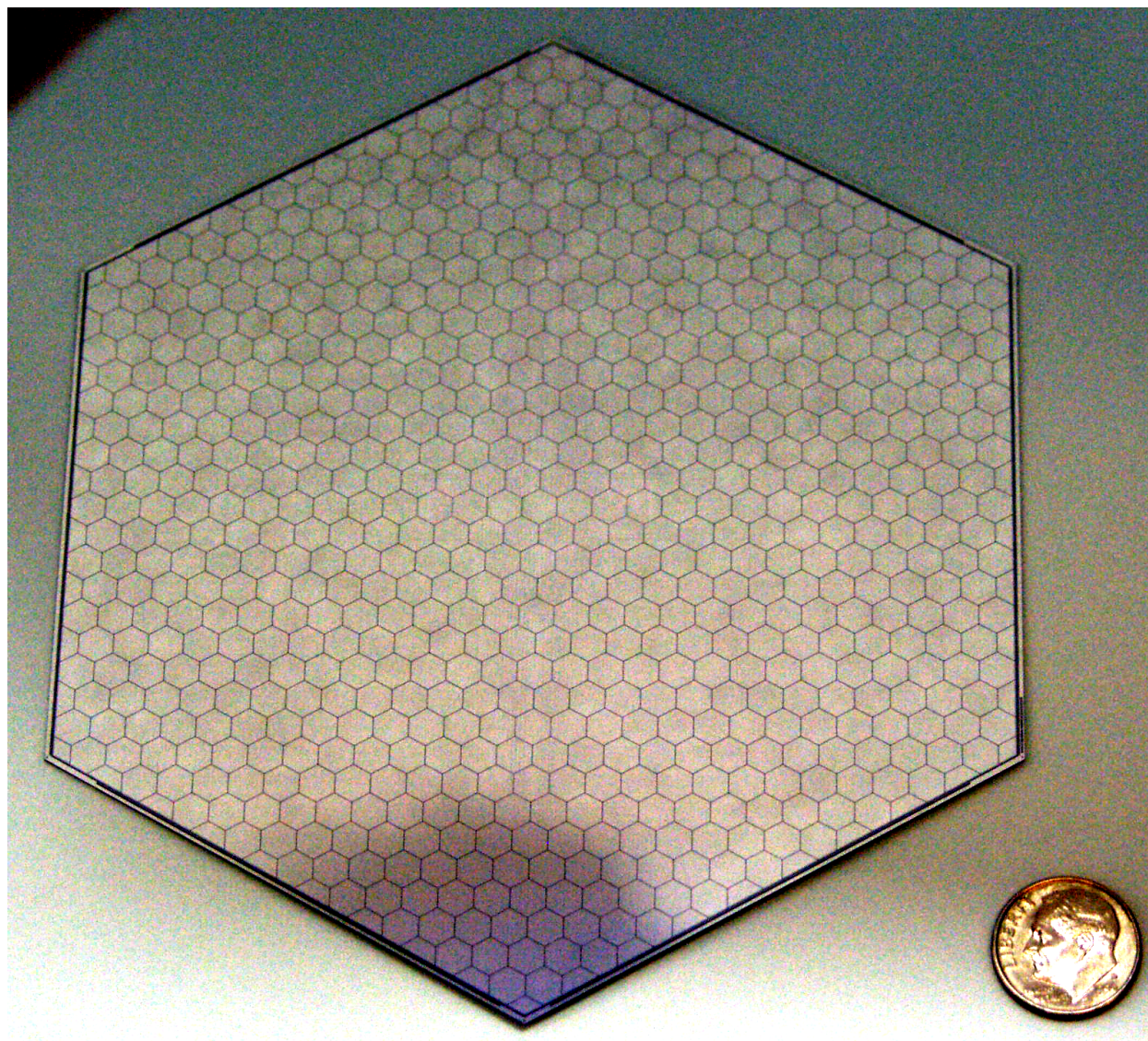


Silicon Detector Design

- DC coupled detectors (avoids bias resistor network)
- Two metal layers
- Keep Si design as simple as possible to reduce cost
- Cross talk looks small with current electronics design
- Trace capacitances (up to 30pF) are bigger than the 5pF pixel capacitance

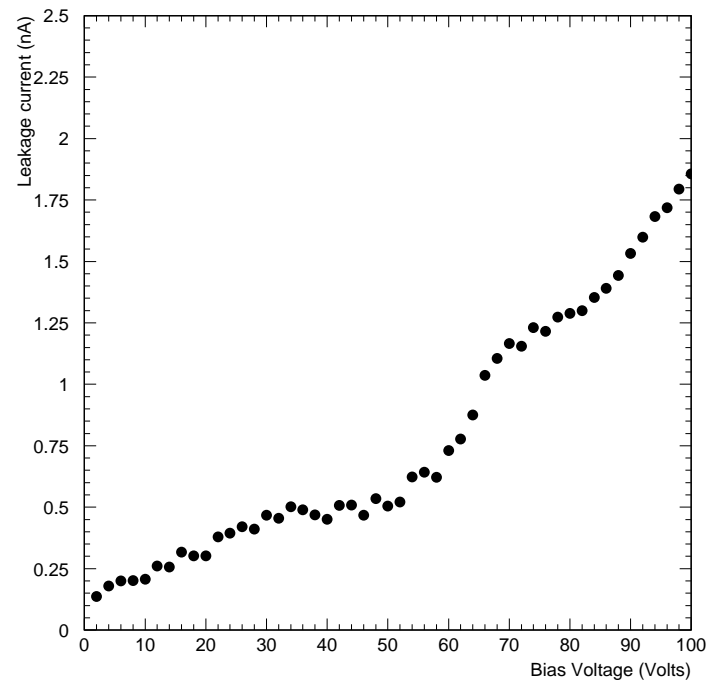


Ten Hamamatsu detectors are in hand



Measurements on Silicon Detector Prototypes

Leakage Current Looks Fine:



(10nA for $1\mu\text{s}$ gives only 250 electrons noise)

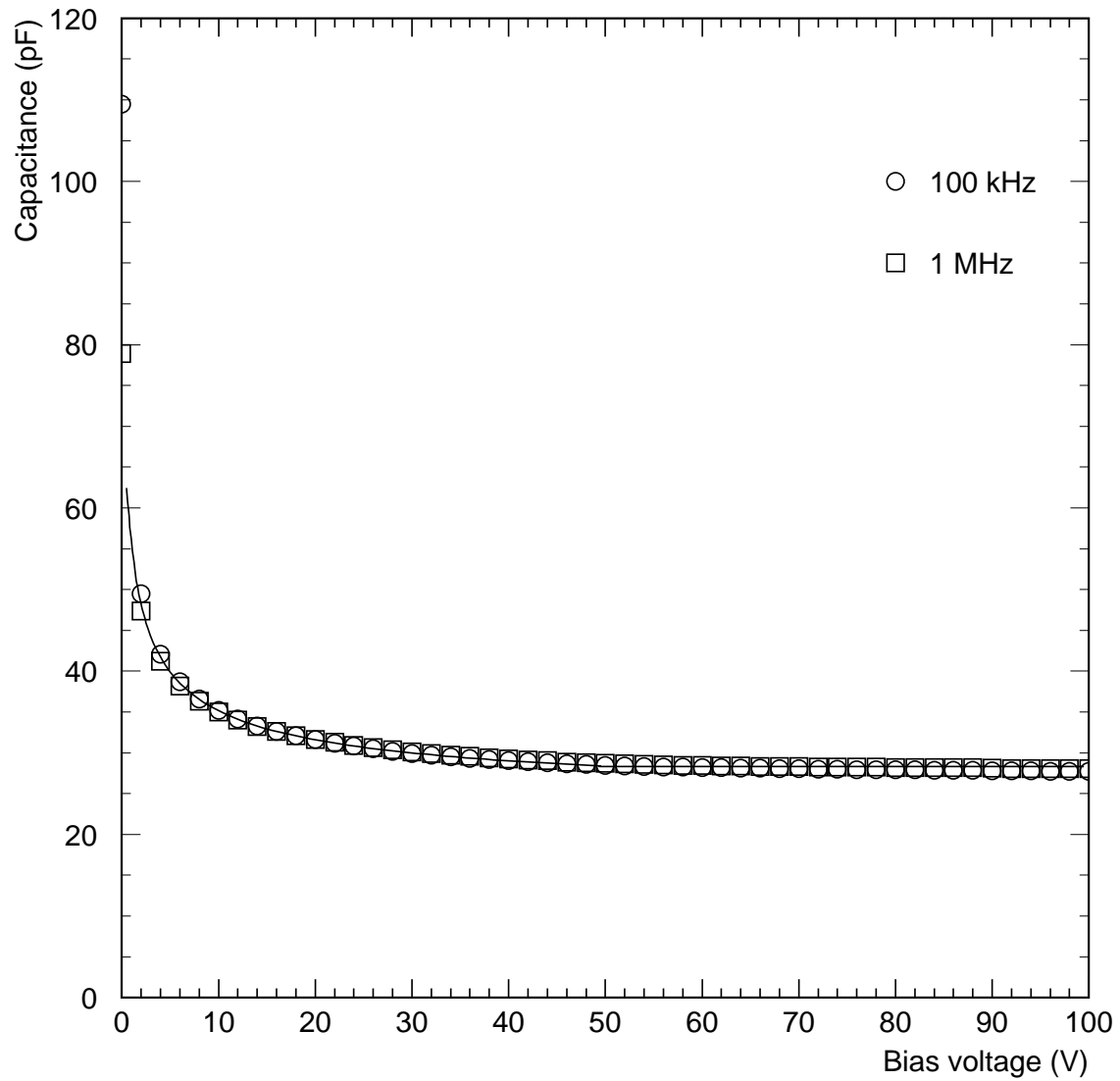
NB: Neighboring pixels are not grounded.

Expected contributions to detector capacitance:

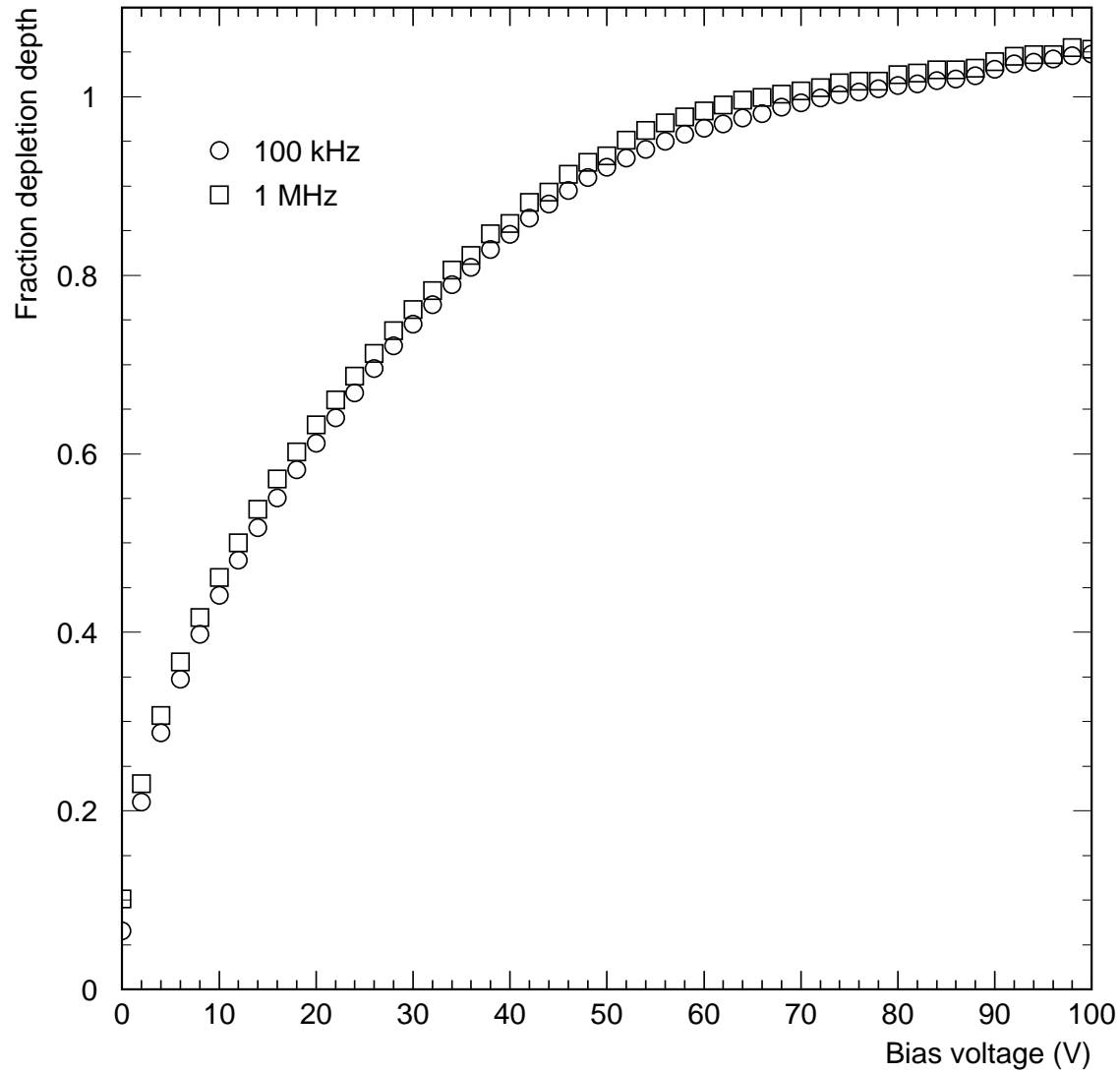
- 5.7pF from pixel capacitance (C_{geom})
- ~ 20 pF for sum of trace capacitance and capacitance from other traces connecting to other pixels. (C_{stray})
- Pixels under the bump-bond array have additional stray capacitance from probing and bonding pads (currently $\simeq 100$ pF)

Expected curves

$$\begin{aligned} C_{tot} &= C_{stray} + C_{geom} \sqrt{\frac{V_{dep} + V_{bi}}{V_{bias} + V_{bi}}} & V_{bias} < V_{dep} \\ C_{tot} &= C_{stray} + C_{geom} & V_{bias} > V_{dep} \end{aligned}$$



Typical CV curve as measured in lab



Relative depletion depth as a function of voltage.

Mean stray capacitance measurement obtained from a fit to the CV curve:

Expected	100kHz	1 MHz
23.0 ± 0.2 pF	21 ± 1 pF	22 ± 1 pF

⇒ **Measurement agrees with expectation for $0.9 \mu\text{m}$ thick oxide and $6\mu\text{m}$ wide traces (3.1 pF/cm).**

Series resistance for $1\mu\text{m}$ by $6 \mu\text{m}$:

Expected (pure Al)	Measured
$47 \Omega/\text{cm}$	$(57 \pm 2)\Omega/\text{cm}$

⇒ **Measurement slightly larger than nominal**

Impact of Detector Technology on Detector Design

⇒ In a warm machine, exceptional pixels with large capacitance or series resistance lead to degraded time tag measurements

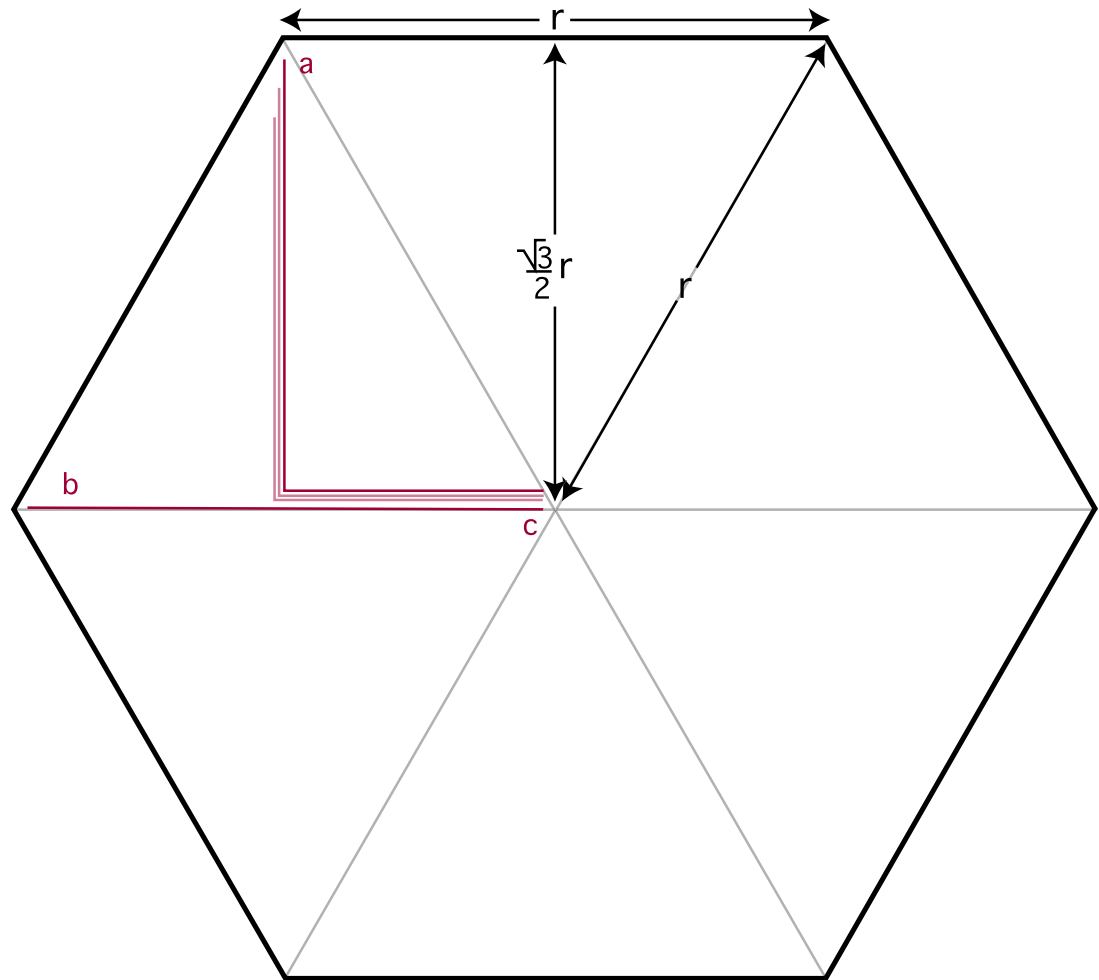
- Small impact on tagging performance since bad channels can be de-weighted in determining the average time of a track

⇒ In a cold machine, exceptional pixels with large capacitance or series resistance lead to a higher rate of noise events in buffers

- Could lead to inefficiency late in the bunch train due to buffer overflow

Location of high resistance and capacitance pixels

- a.) Longest trace ~ 10 cm
- b.) Radial trace ~ 7 cm
- c.) Congested area near bump bond array



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- For areas a and b fundamental limit to noise is given by (for e.g. correlated double sampling)

$$ENC_{R_s} \sim C_{tot} \sqrt{4 \frac{KT}{q_e^2} R_s \frac{1}{2\tau}}$$

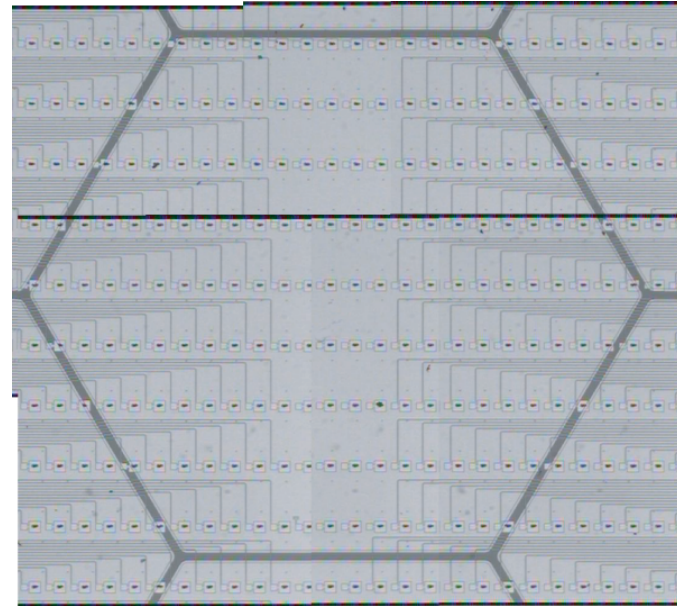
where R_s is the series resistance, C_d and τ is the shaping time of the electronics.

- For $\tau = 1\mu s$, $R_s = 580\ \Omega$ and $C_{tot} = 40\text{ pF}$ this gives ~ 600 electrons noise, which is not really a problem.
- We can slightly improve noise performance by decreasing the trace width, perhaps by a factor of 2, i.e.

$$ENC_{R_s} \propto \sqrt{w}$$

where w is the trace width.

- In region c , near the bump bonding array, we will have a large number of traces crossing a pixel. No series resistance, but amplifier FET noise similar:



Possible ways to decrease capacitance in region c :

- Move probing pads on to pixels.
- Decrease trace width in area near central pixels, here

$$ENC_{amp} \propto w$$

- Use a long skinny chip (e.g. $100 \mu\text{m} \times 600 \mu\text{m}$ grid)

After these three measures, worst case capacitance is ~ 70 pF.

Other more radical alternatives

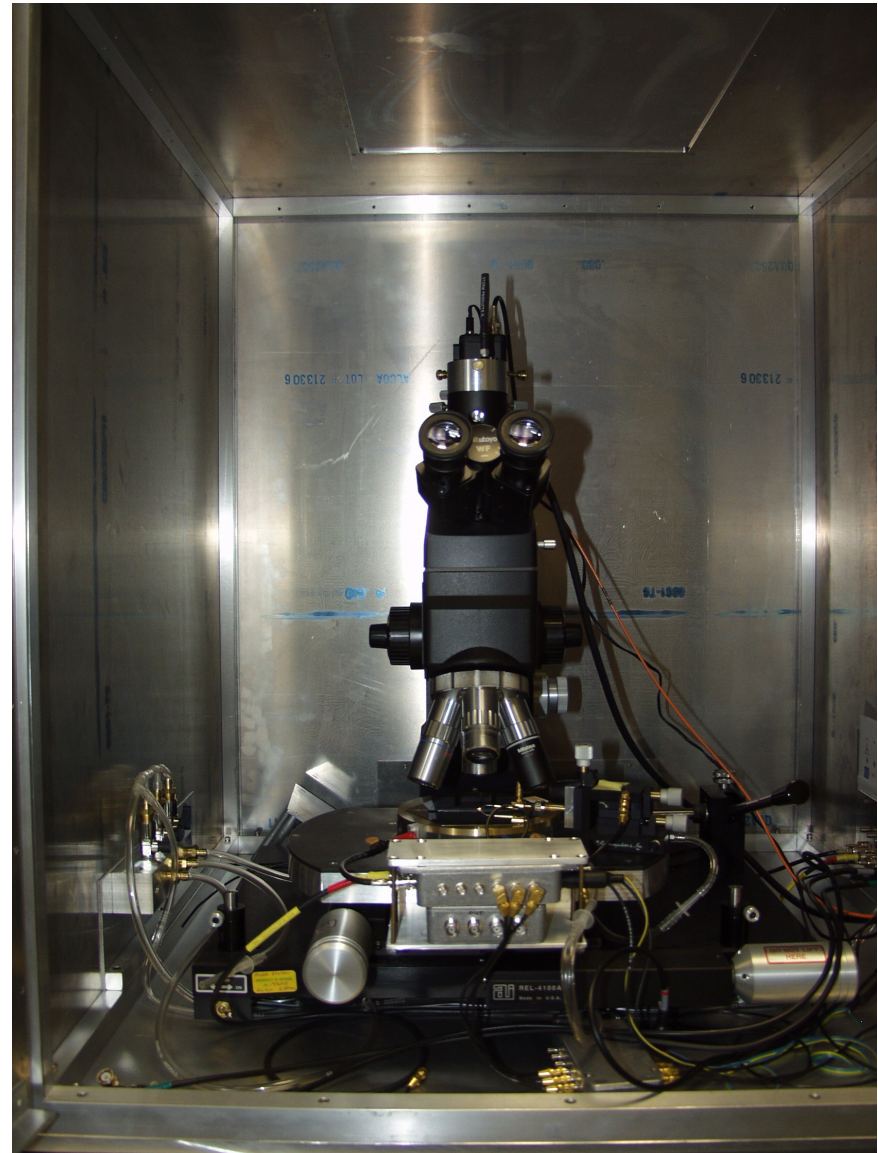
- Polyimide (kapton) can be used instead of SiO_2 as insulator for traces
- Oxide thickness to $5\mu\text{m}$ possible.
- Minimum trace with probably $10\mu\text{m}$
- Could reduce stray capacitances by a factor of 2 or more

Hamamatsu does not currently provide metal-on-polyimide products, but we could increase the thickness of the wafer and the SiO_2 .

SINTEF (Norway) may be producing detectors based on 6 inch wafers with metal-on-polyimide within the next year. (Possible collaboration with Brookhaven to produce masks.)

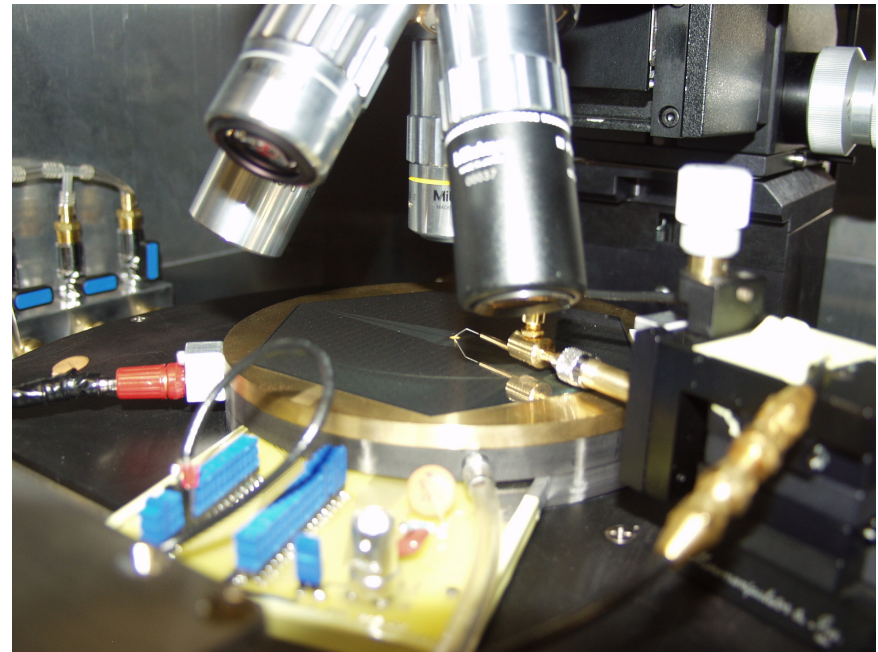
Test Setup for Cosmics, Sources and Laser

- Modified probe station, allows laser to be target on entire detector
- IR microscope objective used to focus laser to $\sim 10\ \mu\text{m}$ spot
- Bias applied to backside of detector using insulated chuck

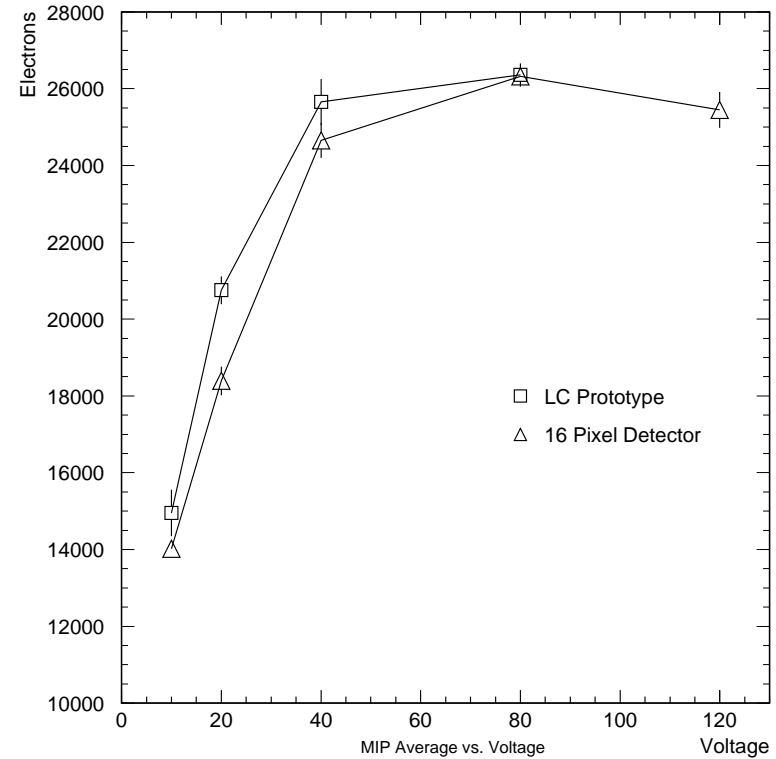
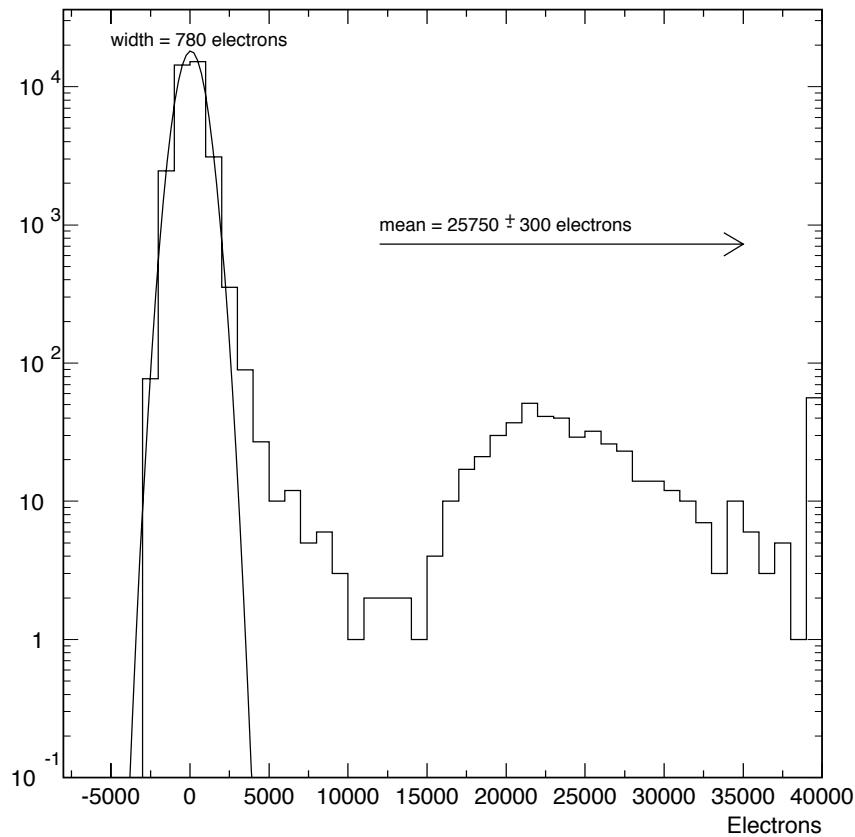
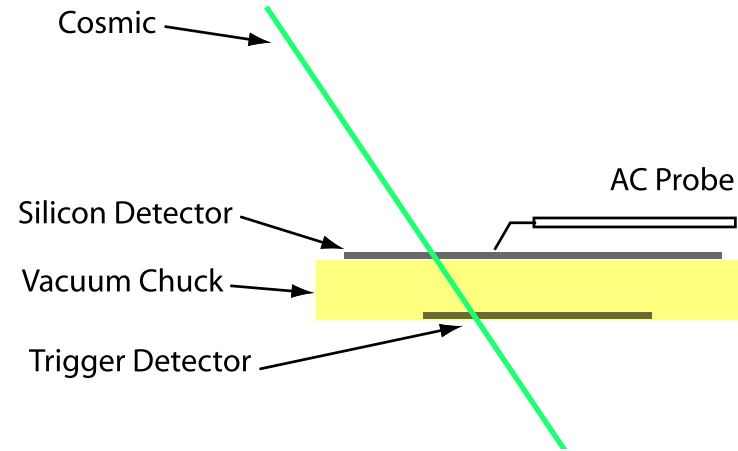


Test Setup – detector probing

- Contact made to test pads on bump bonding array using an AC probe
- Cables add ~ 20 pF of additional capacitance, but noise performance is somewhat better than readout chip
- Use AMPTEK 250F preamp, shapers with $\tau \simeq 1\mu\text{s}$ and a digitizing oscilloscope to mockup expected electronics
- PC board with 1 cm \times 1 cm silicon pad detector used for cosmic trigger visible under chuck

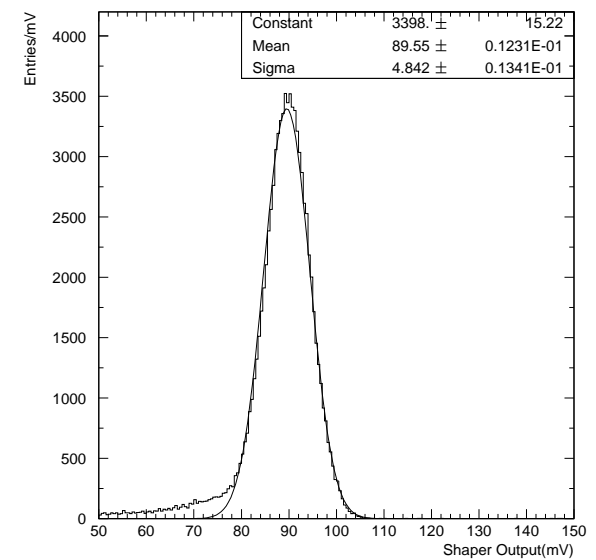
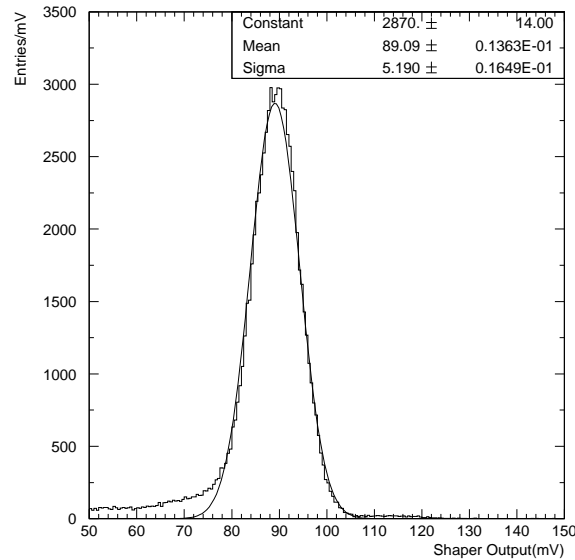
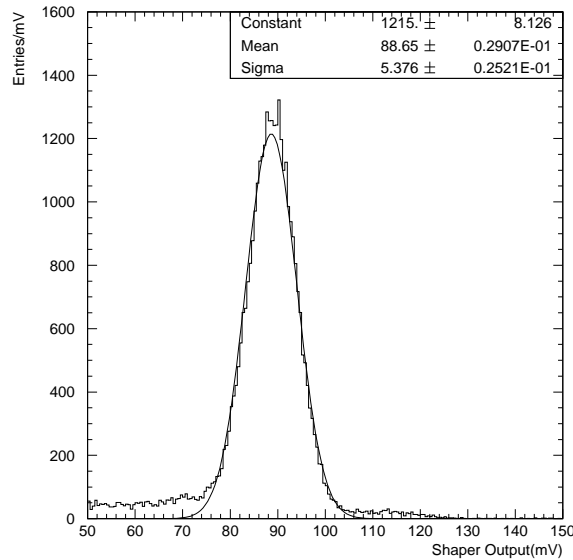


Response of detectors to Cosmics
 (Single 5mm pixel)
 Simulate LC electronics
 (noise somewhat better)



Errors do not include $\sim 10\%$ calibration uncertainty (no source calibration)

Response of Detectors to 60KeV Gamma's from Am²⁴¹



Possible $\sim 1\%$ wafer-wafer calibration?

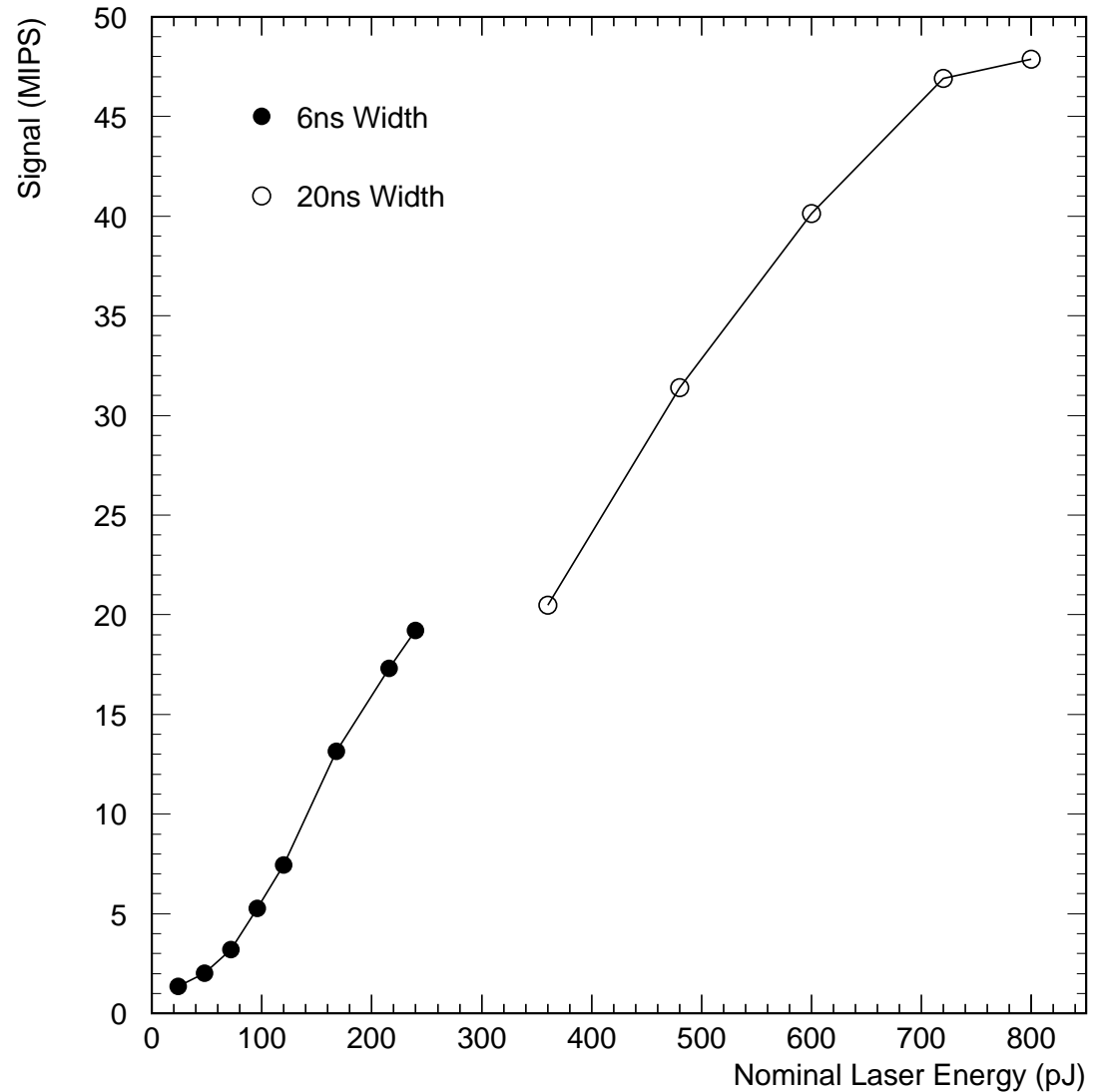
Width of distributions corresponds to ~ 970 electrons noise. Pixels under test are on outer edge of wafer – includes larger series resistance contribution than cosmic data.

Laser Studies

$\lambda = 1064 \text{ nm}$

IR penetrates into wafer

Allows controlled study of large and small pulses



Conclusions

- A narrow gap silicon–tungsten detector for LC physics is attractive
- First round of prototype silicon detectors perform as expected
- Detectors can be produced with workable values of stray capacitance and series resistance
⇒ some changes need for cold design